

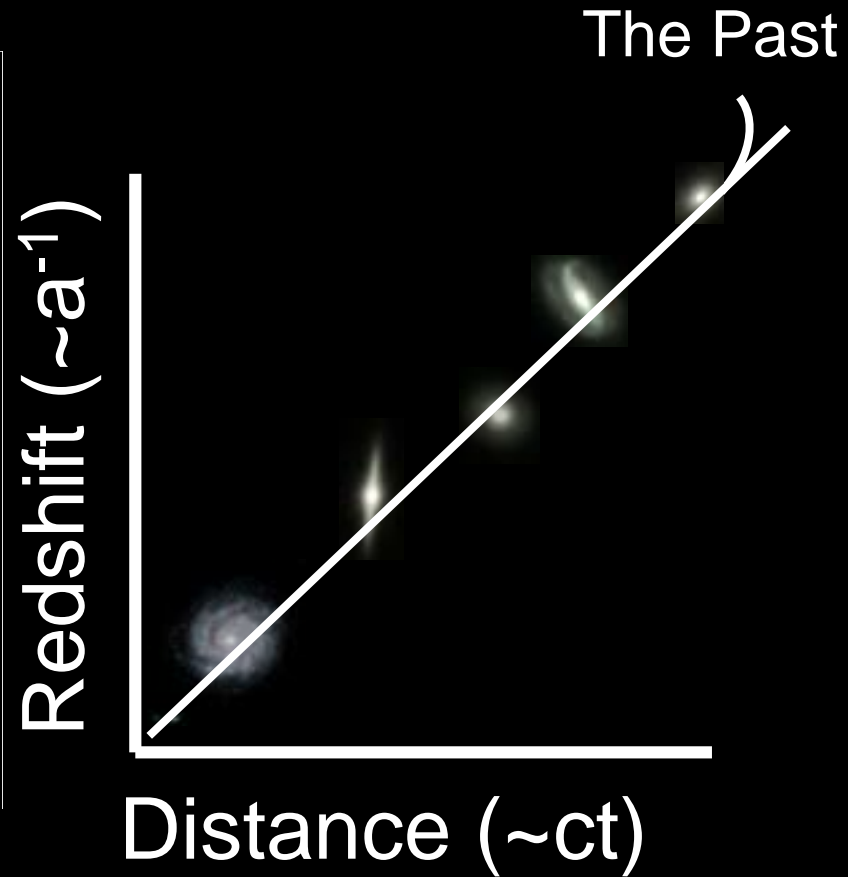
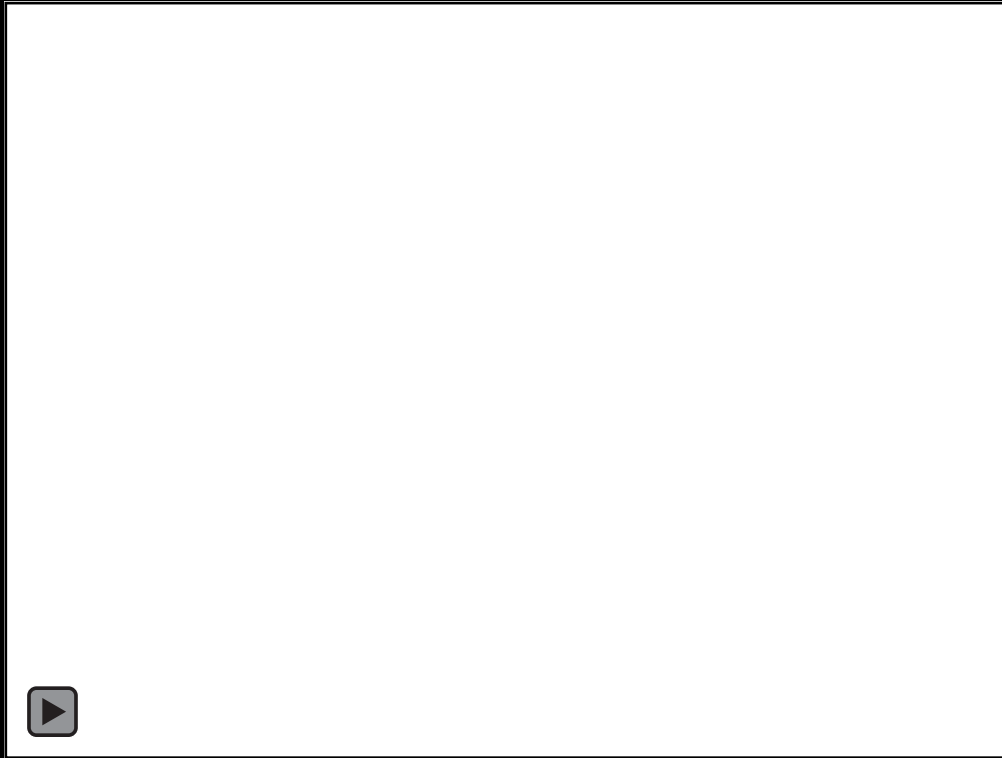
Dr. Adam Riess

Johns Hopkins University  
Space Telescope Science Institute

**COSMIC EXPANSION SEEN BY  
HST,  
“STANDING ON THE SHOULDERS  
OF GIANTS”**



# Expanding Universe reveals Composition, Age, Fate...



Homogeneous, Isotropic + GR  $\rightarrow$   
equation of expansion  $a(t)$ , “scale factor”  
Depends on present state, composition of Universe

Friedmann Equation

$$\left(\frac{\dot{a}}{a}\right)^2 = \frac{8\pi G \rho_M}{3} + \frac{\Lambda}{3} - \frac{k}{a^2}$$

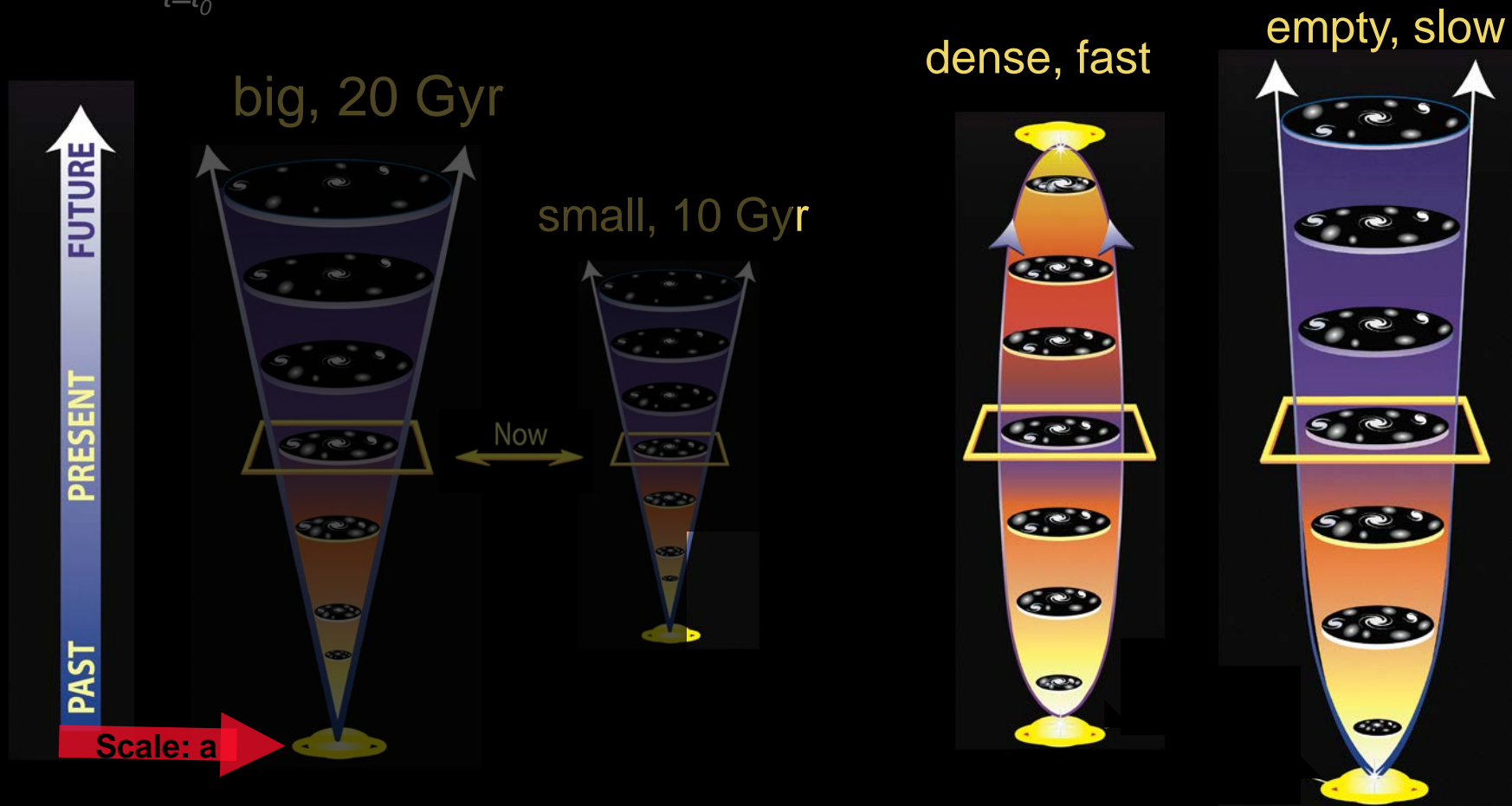
# Before HST: Cosmology was the quest for two numbers

$$H_0 = \left. \frac{\dot{a}}{a} \right|_{t=t_0}$$

Present rate, size, age,  
Key Project of HST!

$$q_0 = \left. \frac{-\ddot{a}}{aH_0^2} \right|_{t=t_0}$$

Deceleration by  $\Omega_M (=2q_0)$ , geometry, fate  
origin, viability of inflation



1990's: Better  $D(z)$  with long range Standard Candles, SN Ia...

# Original Objective of HST #1, Gauging the Universe

## From Lyman Spitzer, 1946, RAND Report

### 1. Extent of the Universe

The 200-inch telescope is designed to push back the frontiers [of] explored space. It is not likely that this instrument will reach to the greatest distance possible. Further measurements with the more powerful instrument envisaged here would help answer the questions whether space is curved, whether the universe is finite or infinite. This instrument would help in particular to resolve individual stars in a distant galaxy and to analyze their spectra, thus identifying particular stars of known absolute magnitude and in this way determining accurately the distance to the galaxy. At present the distances of most galaxies are known only very approximately.



## Colgate, S, ApJ, 1979

“Supernovae type I can perhaps be found to  $z=1$  using the Space Telescope...to accurately determine  $q_0$ , the cosmological constant”



# SN Ia Hubble Diagram; $q_0$ , Accelerating Universe, Dark Energy!

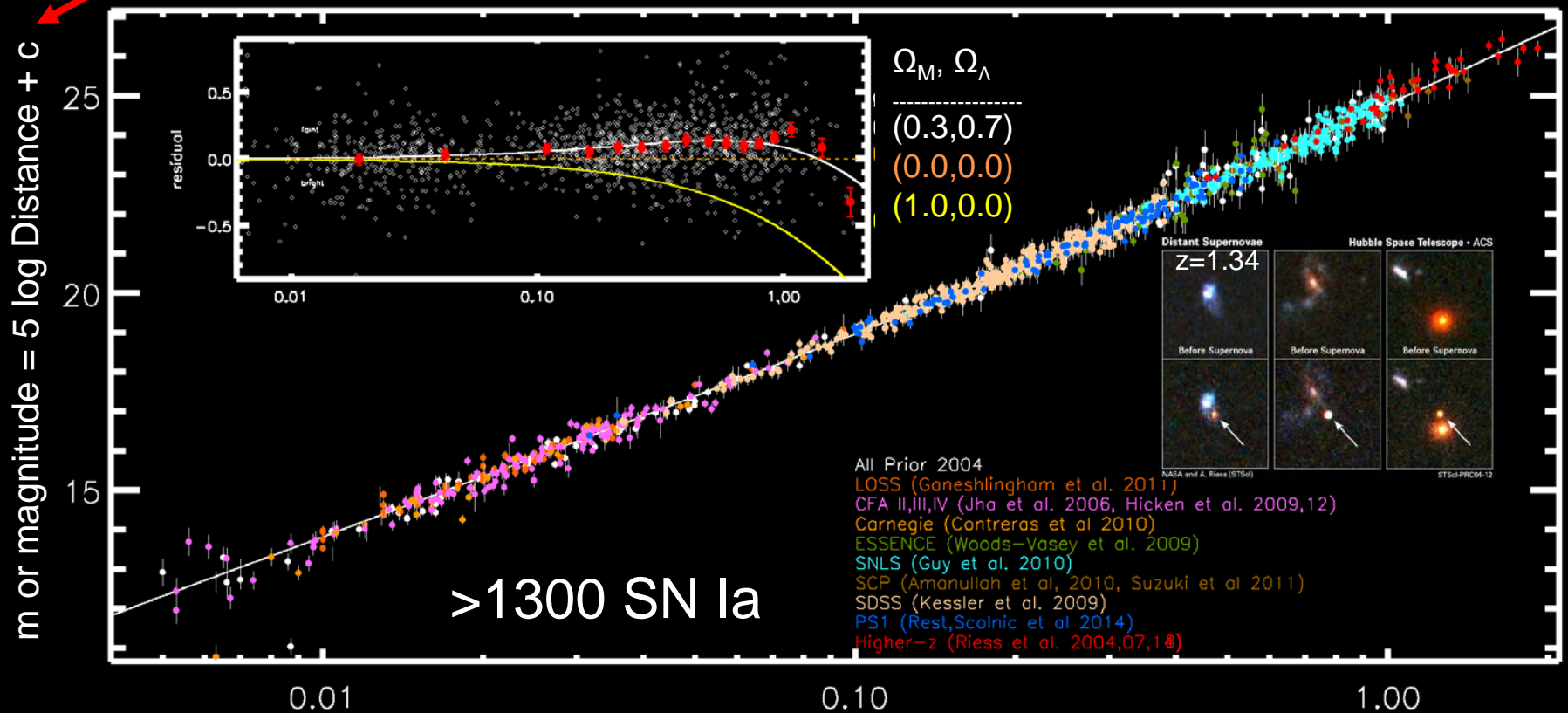
## HST's Unique Contributions:

1996-2001: WFPC2, follow-up some grnd SNe Ia at  $z < 1$ , best data

2001-2007: ACS+NIC2: find SN Ia at  $z > 1$ , confirming "turn-over"

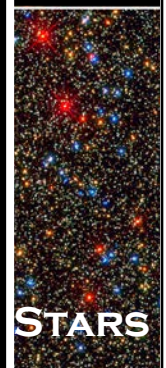
2007-present: WFC3: characterizing  $w(z)$ , looking for unexpected

Independent of absolute distance,  $H_0$  (2001: KP to 10%)



# 2010's: "End-to-end" test for $\Lambda$ CDM Predict and Measure $H_0$

Standard



$$\frac{\ddot{a}}{a} = -\frac{4\pi G}{3} \left( \rho + \frac{3p}{c^2} \right) + \frac{\Lambda c^2}{3}$$

Planck Predicted,  $H_0=67.4\pm 0.5$  km/s/Mpc

# After KP: Direct Measurement of $H_0$ to percent precision

## The $SH_0ES$ Project (2005)

(Supernovae,  $H_0$  for the dark energy Equation of State)

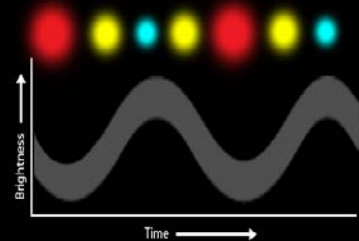
A. Riess, L. Macri, S. Casertano, D. Scolnic, A. Filippenko, W. Yuan, S. Hoffman, et al

Measure  $H_0$  to percent precision purely empirically by:

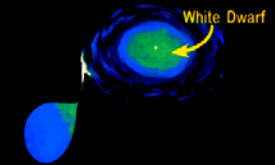
- A clean, simple ladder: **Geometry  $\rightarrow$  Cepheids  $\rightarrow$  SNe Ia**



Pulsating Stars,  
 $10^5 L_{\odot}$ , P-L relation



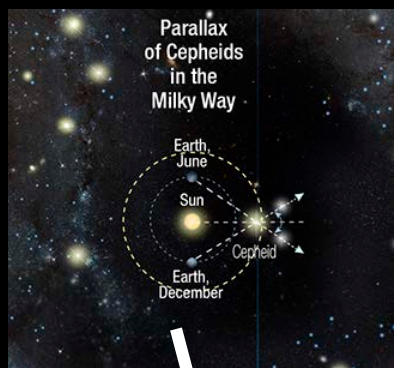
Exploding Stars,  
 $10^9 L_{\odot}$ ,  $\sigma \sim 5\%$



An explosion resulting from the thermonuclear detonation of a White Dwarf Star.

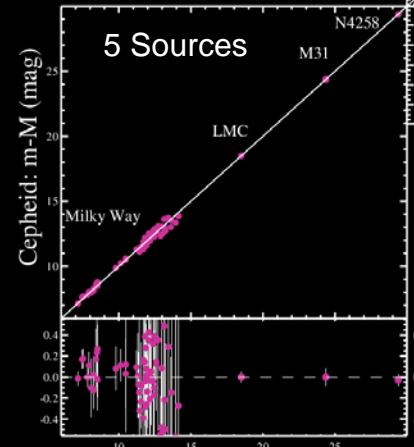
- Reducing systematic error with better data, better collection
- Thorough propagation of statistical and systematic errors

# The Hubble Constant in 3 Steps: Present Data



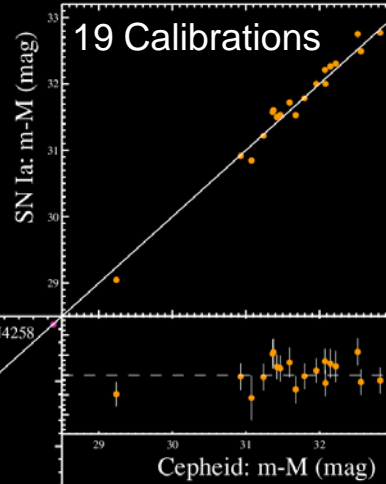
1

Geometry → Cepheids



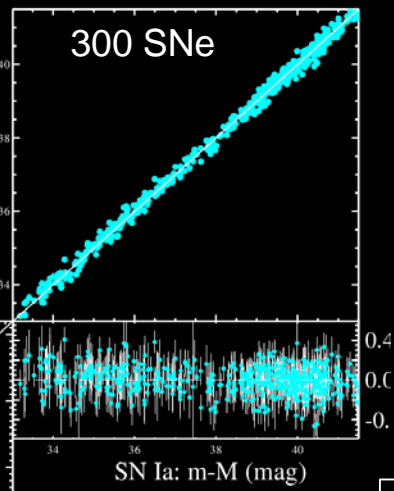
Geometry:  $5 \log D$  [Mpc] + 25

Cepheids → Type Ia Supernovae



HST

Type Ia Supernovae → redshift(z)

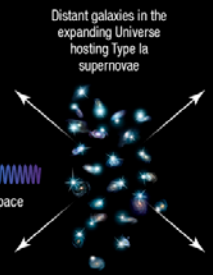


2

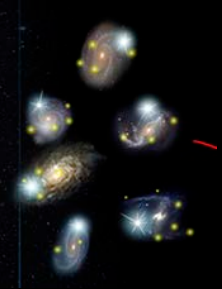
$H_0 = 74.03 \pm 1.42$ ,  
 $\text{Km s}^{-1} \text{Mpc}^{-1}$   
 (Riess et al. 2019)

3

Light redshifted (stretched) by expansion of space



Galaxies hosting Cepheids and Type Ia supernovae



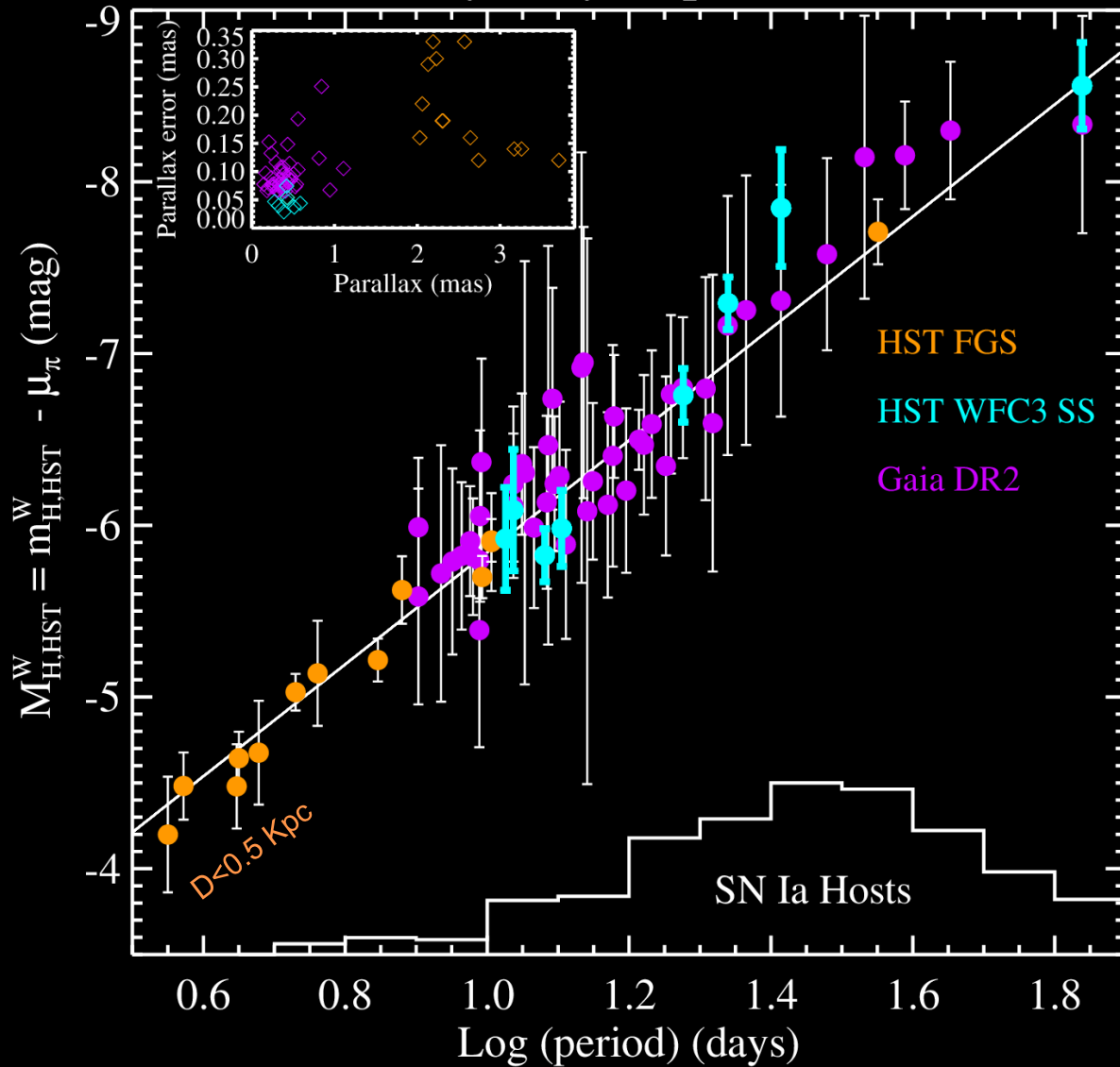
1.9% total uncertainty

$\neq \text{CMB} + \Lambda\text{CDM}!$



# Milky Way Cepheid P-L Relation, Now w/ HST photometry, Long Periods

## Milky Way PL Relation



Final Gaia Parallaxes  
+ HST Photometry  $\rightarrow$   
 $H_0 \sim 0.4\%$

} with 3 band  
HST photometry  
and  
Periods > 10 days  
both matching  
Cepheids HST sees  
in SN Ia hosts

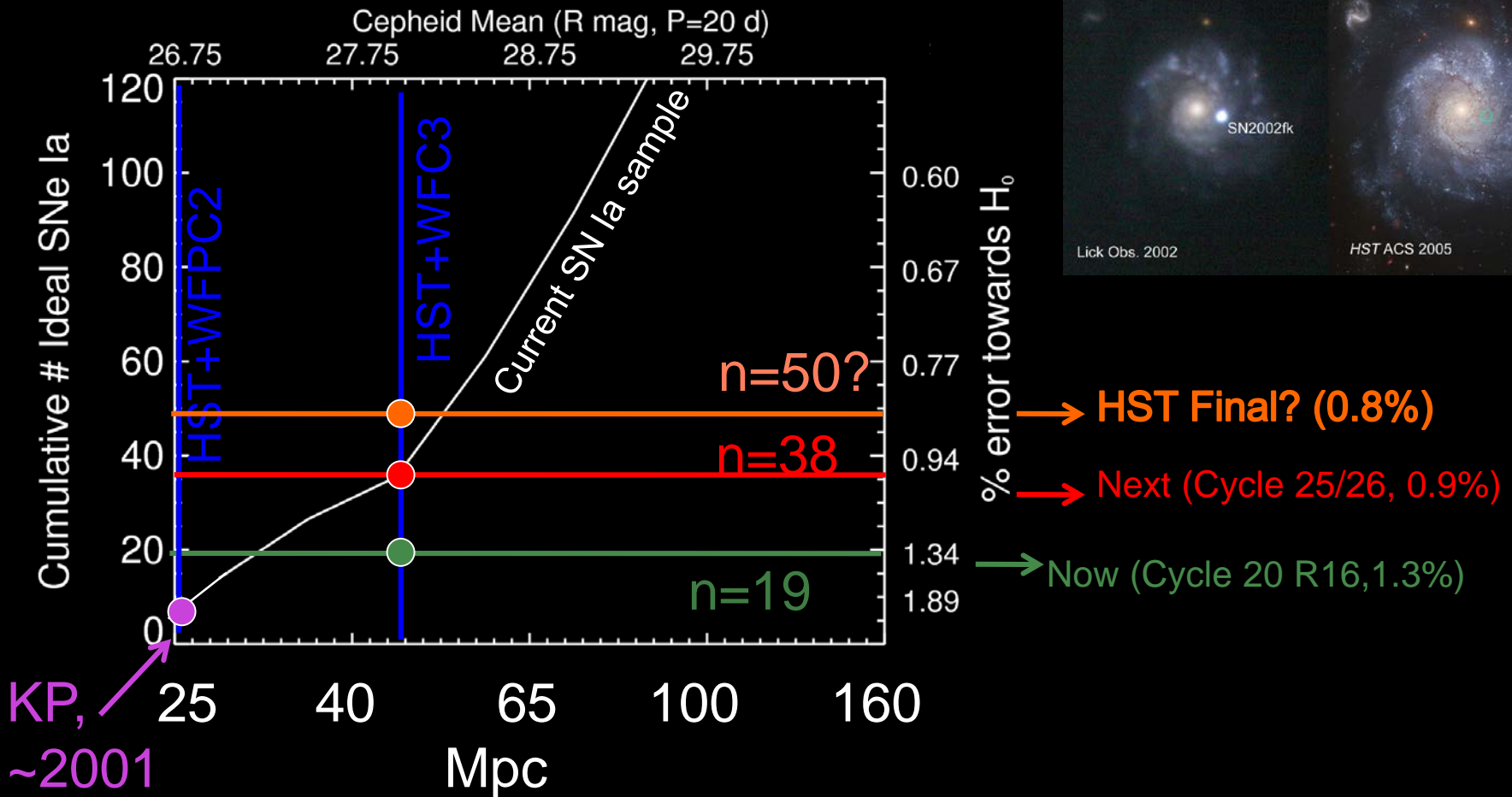
# Robust? Five Sources of Cepheid Geometric Calibration

	<b>Independent Geometric Source</b>	$\sigma$	$H_0$
	NGC 4258 H <sub>2</sub> O Masers: Humphreys et al 2013, Riess et al 2016	2.6%	72.3
<b>NEWER</b>	LMC 20 Late Detached Eclipsing Binaries: Pietzrynski et al. 2019 +70 HST LMC Cepheids Riess et al (2019)	1.3%	74.2
	Milky Way 10 HST FGS Short P Parallaxes: Benedict et al. 2007 --also Hipparcos (Van Ileeuwen et al 2007)	2.2%	76.2
<b>NEW</b>	Milky Way 8 HST WFC3 SS Long P Parallaxes: Riess et al. 2018	3.3%	75.7
<b>NEW</b>	Milky Way 50 Gaia+HST, Long P Parallaxes: Riess et al. 2018	3.3%	73.7

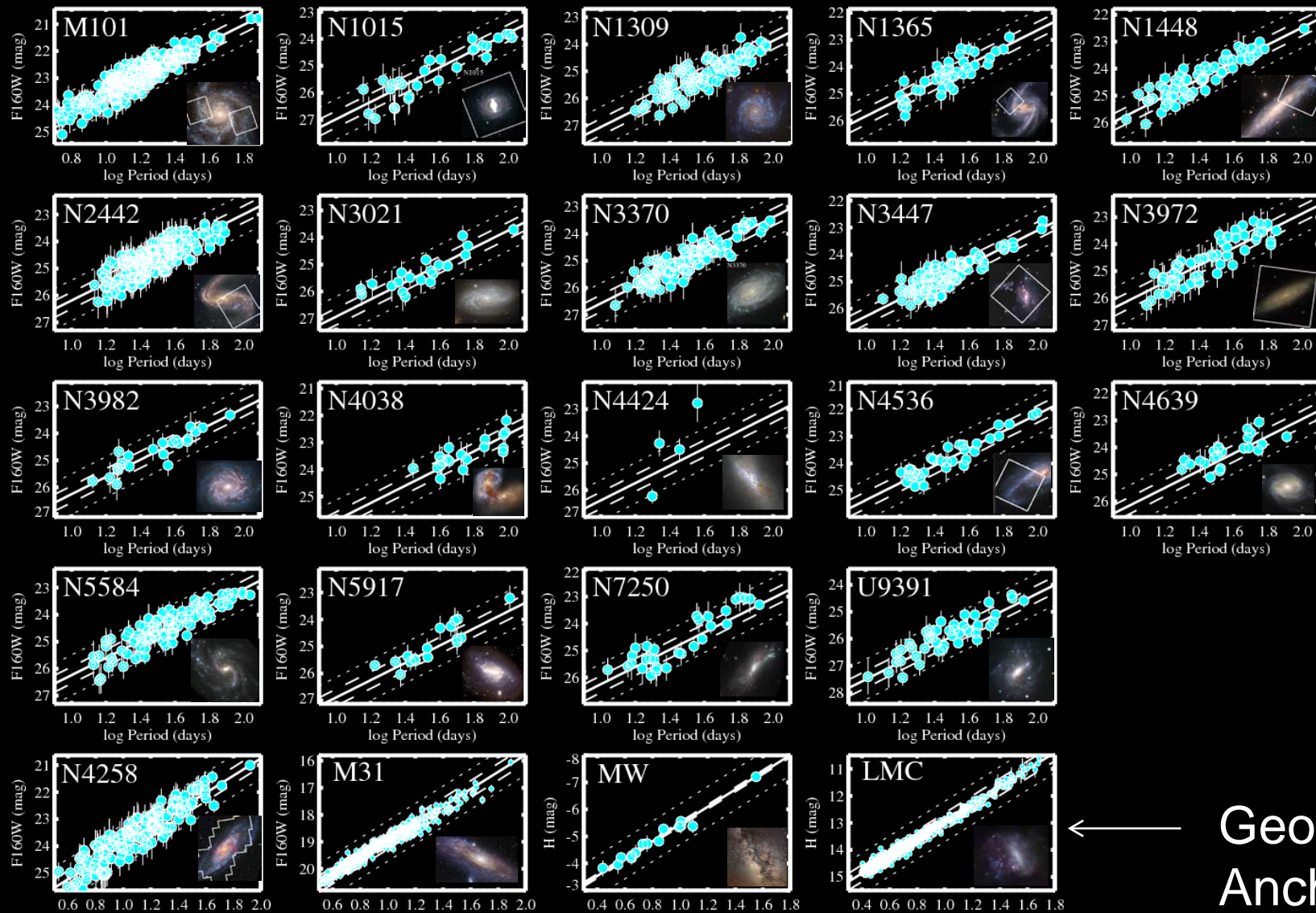
Consistent Results ( $1.3\sigma$ ), *Independent Systematics*

# Step 2: Cepheids to Type Ia Supernovae

This is the  $H_0$ -Limiting Step: Number of SN Ia in Cepheid Range



# Cepheid V,I,H band Period-Luminosity Relationships: 19 hosts, 3 anchors



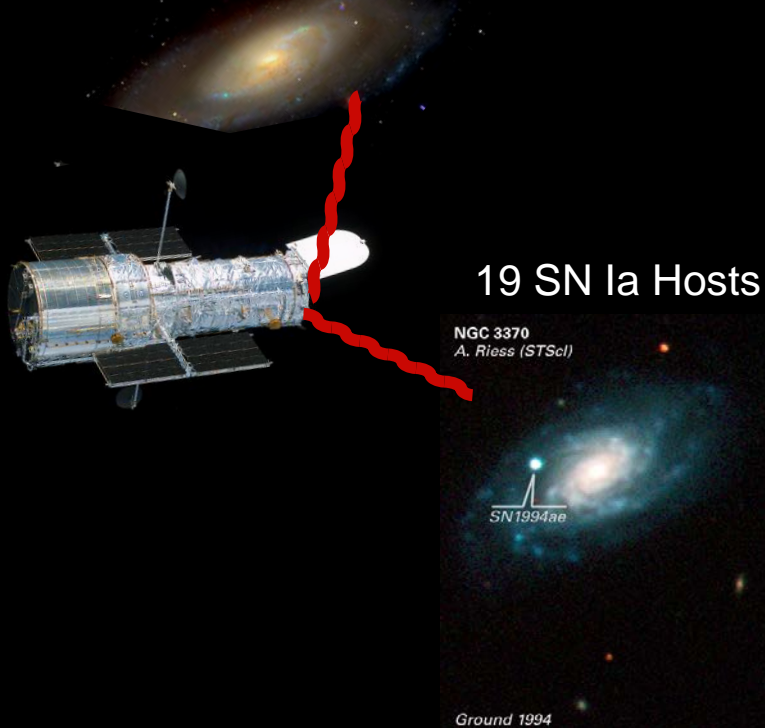
S  
N  
H  
O  
S  
T  
S

← Geometric Anchors

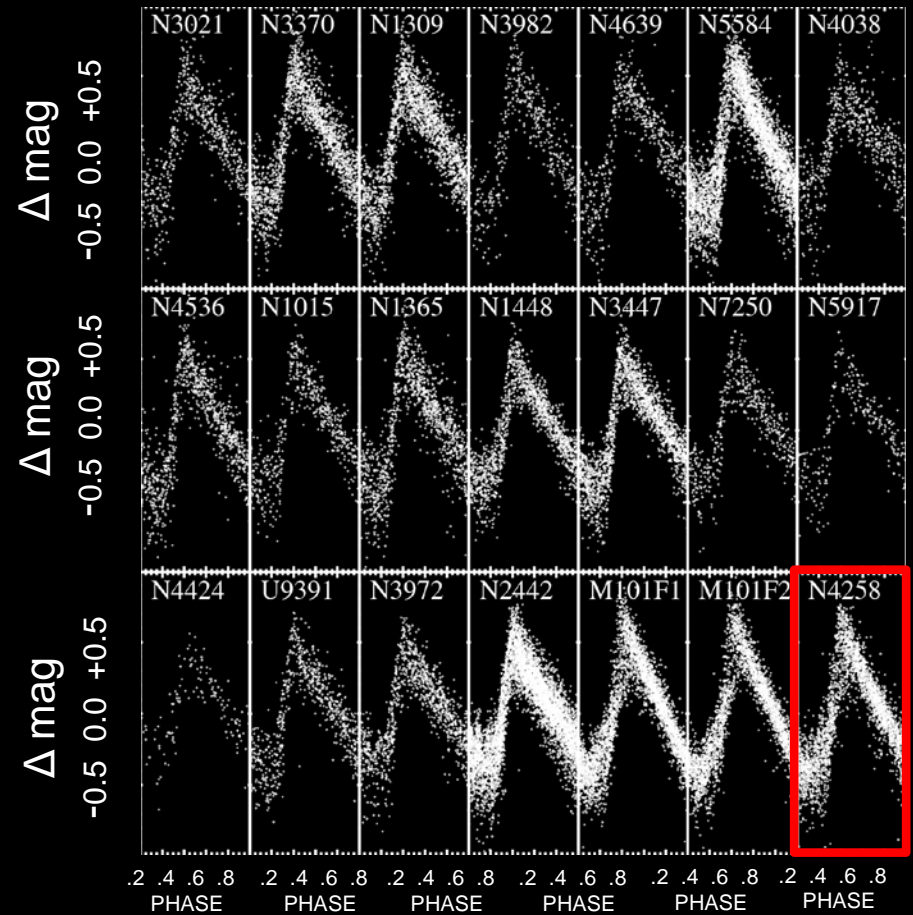
# Lower Systematics from *Differential* Flux Measurements

We reduce systematic errors by measuring all Cepheids with same instrument, filters, similar metallicity, period range, we correct for crowding and dust statistically

ANCHORS: NGC 4258 (and now MW, LMC) geometric distance



Cepheid composite LC's, >2400



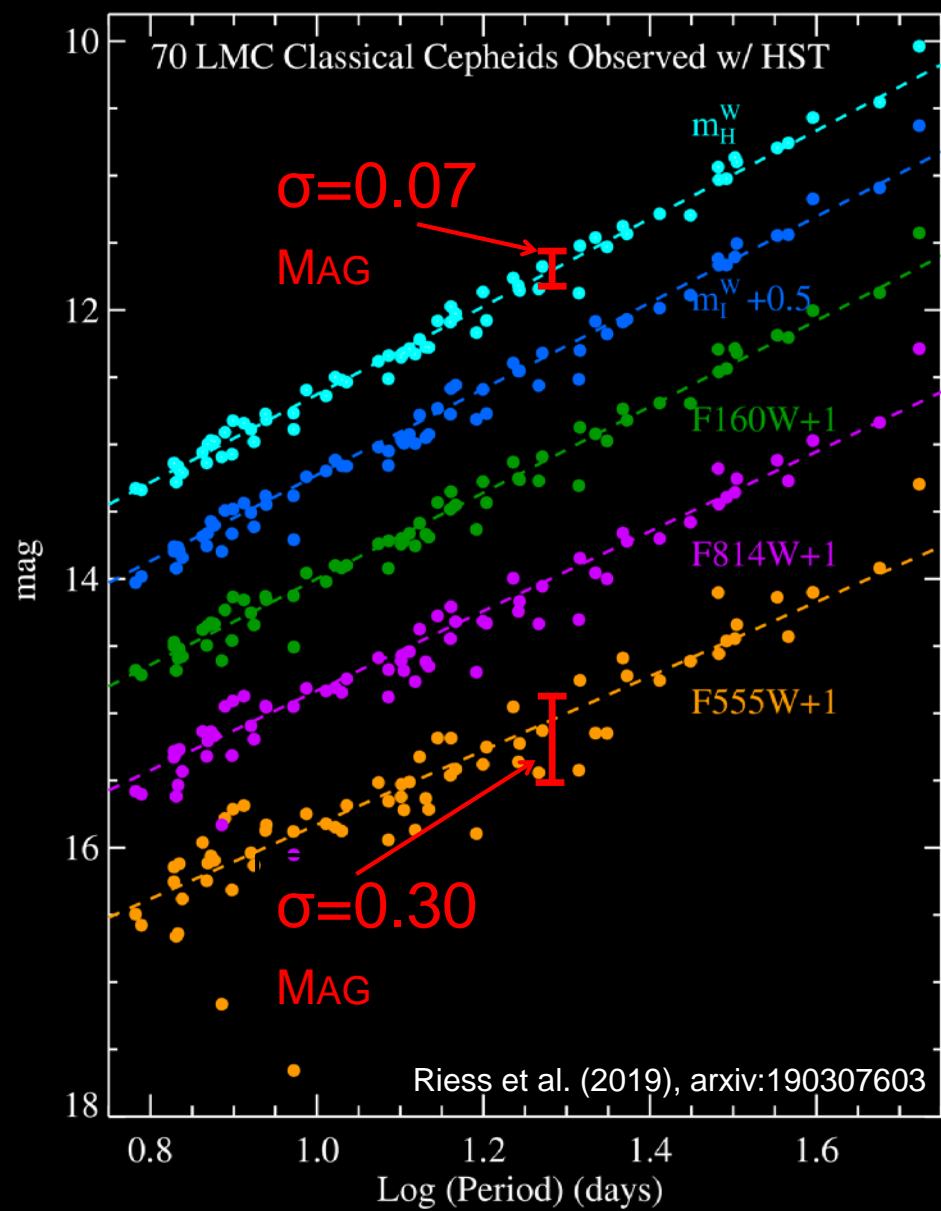
# Lowering Systematics: Near-IR Cepheid Observations + HST, Now in LMC!

-Negligible sensitivity to metallicity in NIR (F160W)

-Dependence on reddening laws 6x smaller than optical

We use F160W-band as primary +F555W,F814W

Key Project used F555W and F814W



Dereddened:  
F160W-0.386(F555W-F814W)

1.6 m

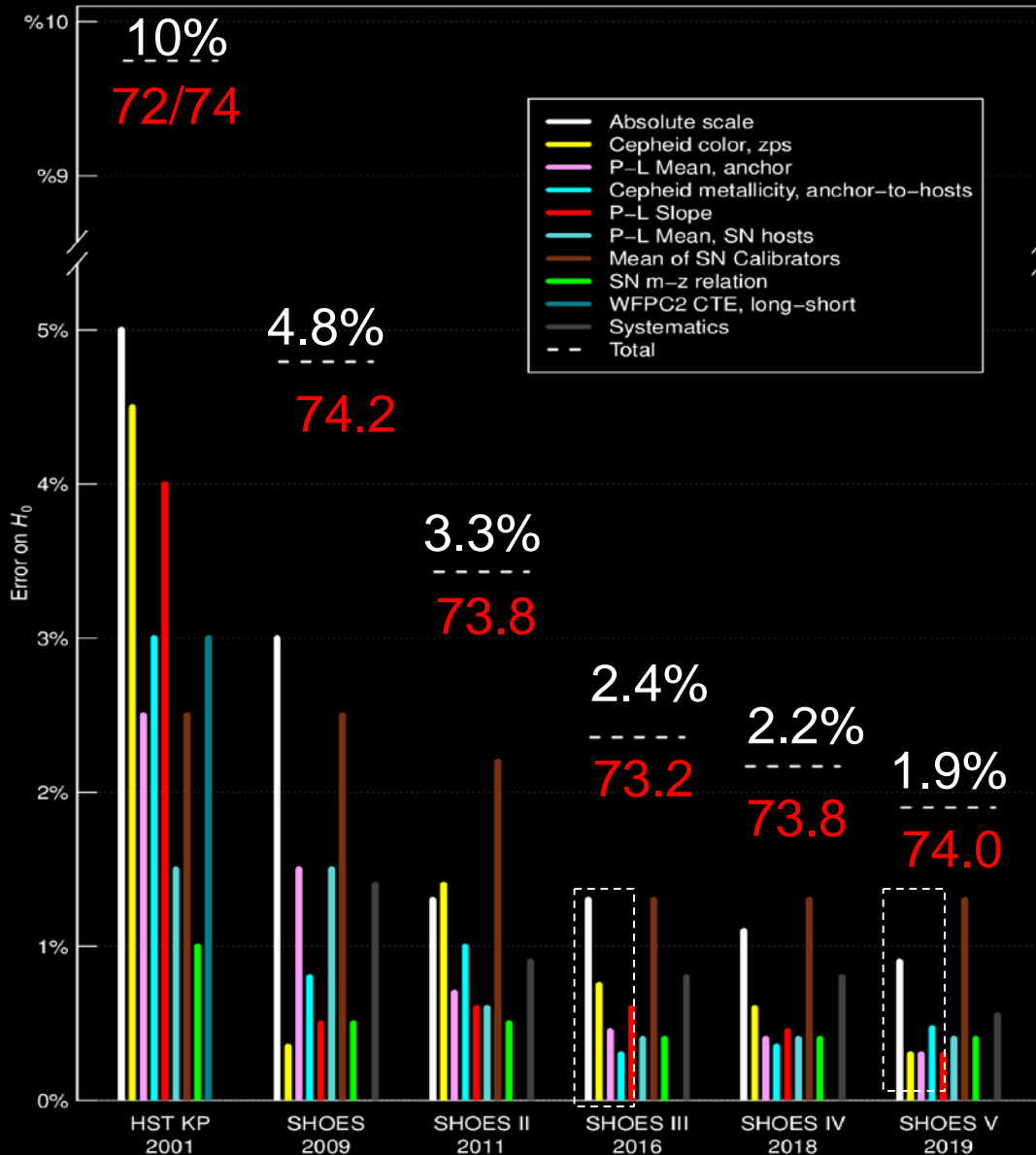
0.8 m

0.55 m



Leavitt

# HST Distance Ladder Error Budgets for $H_0$ (w/ SN+Cepheids) 2001-2019

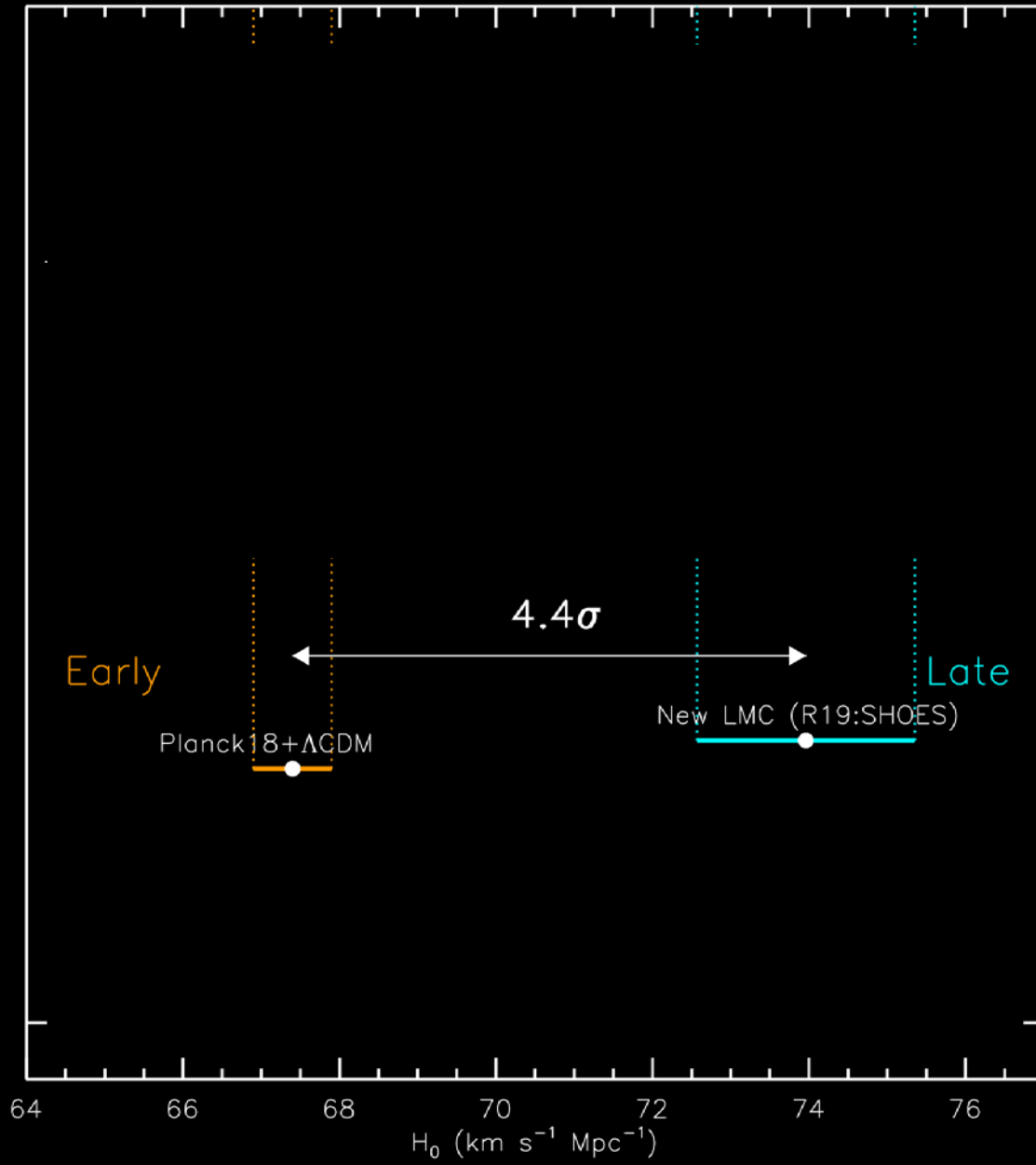


Main improvements  
 Since 2016:  
 Anchors—MW  
 parallaxes, LMC  
 DEB distance,  
 matched Cepheid  
 photometry, WFC3  
 CRNL

How does this compare to the CMB measurements?

# $H_0$ : Measured Late vs. Predicted from Early Universe

NEW  
PHYSICS





# Breakthroughs When Local $H_0$ was too high. This time?

1930-1950:

$H_0 > 300 \text{ km s}^{-1} \text{ Mpc}^{-1} \rightarrow t_0 \sim \text{Gyr} \ll \text{age of Earth}$

Why? Two populations of stars! Early and late, poor and rich.



1990's\*:

$60 < H_0 < 85 + \Omega_M = 1 \rightarrow t_0 (10 \text{ Gyr}) \ll \text{oldest stars} (14 \text{ Gyr})$

Why? Dark energy!  $\Omega_M \sim 0.3$ ,  $\Omega_\Lambda \sim 0.7$



2010's:

$H_0 = 74 \pm 1.4 \rightarrow 4.4\sigma$  higher than Planck CMB +  $\Lambda$ CDM

What will be discovered ?



\* Internally inconsistent measures of  $H_0$  indicated systematics not new features

# Takeaways

- Universe now appears to be expanding  $\sim 9\%$  ( $\pm 2.2\%$ ) faster than-expected based  $\Lambda$ CDM+Planck CMB
- There are independent checks on each measurement so, either a *conspiracy* of errors or a new feature of LCDM
- We anticipate significant improvements in these measurements in just the next few years which may reveal the cause.
- With additional measurements HST and Gaia can enable a 1% measurement of  $H_0$ , a benchmark for constraining the cosmological model.