

# Cstat: Inference and Goodness-of-fit

Yang Chen

University of Michigan

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# Outline

Literature

Plan

Literature

## Key Papers

- ▶ Cash, W., *Parameter estimation in astronomy through application of the likelihood ratio*, *Astrophysical Journal*, Part 1, vol. 228, Mar. 15, 1979, p. 939-947.
  - ▶ Inference: MLE & confidence intervals
  - ▶ Goodness-of-fit Test:  $\chi^2$  for difference of likelihood ratios – if there exists a hypothesized fixed subset of parameters?
- ▶ Kaastra, J. S. *On the use of C-stat in testing models for X-ray spectra*, *Astronomy & Astrophysics* 605 (2017): A51.
  - ▶ Goodness-of-fit Test: Approximate Gaussian
- ▶ Gaps and problems:
  - ▶ Numbers of bins larger than number of counts for faint sources
  - ▶ Goodness-of-fit tests are approximate
  - ▶ Approximate likelihood with discretization

# Multiple Attempts

- ▶ Asymptotics for C-stat with small counts per bin and large bin count
- ▶ Asymptotic/conservative test of goodness
- ▶ Dynamic bin split and merge
- ▶ Practical implementation with discretized likelihood

## Results: Asymptotic Normality

### Theorem

Let  $\hat{\theta}_n$  be the maximum likelihood estimate. Assume that  $\{s_i(\theta_0)\}_{i \geq 1}$  is bounded from above and  $\text{rank} \left( \frac{\partial \mathbf{s}_{1:n}(\theta)}{\partial \theta} \Big|_{\theta=\theta_0} \right) = d$ .

1. Assume that for all  $\theta \in \Theta$ ,  $\sum_{i=1}^n [\log s_i(\theta)]^2 = O(n^{1-\alpha})$  for some  $\alpha > 0$ . Then  $\hat{\theta}_n \rightarrow \theta_0$  almost surely as  $n \rightarrow \infty$ .

Furthermore,

$$I_n(\theta_0)^{-1} \left[ - \frac{\partial^2 \log L_n(\theta)}{\partial \theta \partial \theta^\top} \Big|_{\theta=\theta_0} \right] \xrightarrow{P} \mathbf{1} \quad \text{as } n \rightarrow \infty.$$

2. Assume that (a) for any  $\theta$  in a small neighborhood of  $\theta_0$ , each  $s_i(\theta)$  is second order continuously differentiable and  $[\log s_i(\theta)]''$  is uniformly bounded by a finite constant, (b)  $[\log n]^2 / I_n(\theta_0) \rightarrow 0$  as  $n \rightarrow \infty$ ; then  $\sqrt{I_n(\theta_0)} (\hat{\theta}_n - \theta_0) \xrightarrow{D} \mathcal{N}(0, I_d)$  as  $n \rightarrow \infty$ .

## Results: C-stat Property

### Lemma

For any  $n$ ,  $-C_n(\hat{\theta}_n) + C_n(\theta_0) = \text{LR}_n^*$ , where  $\text{LR}_n^*$  is given by

$$\begin{aligned}\text{LR}_n^* &= -2 \log \frac{L(s_1(\theta_0), \dots, s_n(\theta_0) | N_1, \dots, N_n)}{L(s_1(\hat{\theta}_n), \dots, s_n(\hat{\theta}_n) | N_1, \dots, N_n)} \\ &= 2 \sum_{i=1}^n \left[ N_i \log s_i(\hat{\theta}_n) - N_i \log s_i(\theta_0) + s_i(\theta_0) - s_i(\hat{\theta}_n) \right],\end{aligned}$$

which is the likelihood ratio statistics for testing the null hypothesis  $H_0 : \theta = \theta_0$  versus the alternative

$H_1 : \{s_i(\theta), 1 \leq i \leq n\} \in \mathcal{S}$ . As  $n \rightarrow \infty$ ,  $\text{LR}_n^* \xrightarrow{\mathcal{D}} \chi_d^2$ .

## Results: Binning Impacts

### Theorem

*Performing finer partitions does not decrease the Fisher information. In fact, with any finer partition, the Fisher information increases unless in the following situation.*

- ▶ *there exists  $s_j^*(\theta)$ ,  $1 \leq j \leq M$  and a partition of  $\{1, \dots, n\}$ , denoted by  $\{\sigma_1, \dots, \sigma_M\}$ , such that for any  $1 \leq i \leq n$ , there exists  $1 \leq j \leq M$  such that  $i \in \sigma_j$  and  $s_i(\theta) = c_{ij}s_j^*(\theta)$  for some constant  $c_{ij}$*