





# Investigating galaxy evolution with deep learning

#### Marc Huertas-Company

Instituto de Astrofísica de Canarias Observatoire de Paris Université Paris Diderot Institut Universitaire de France

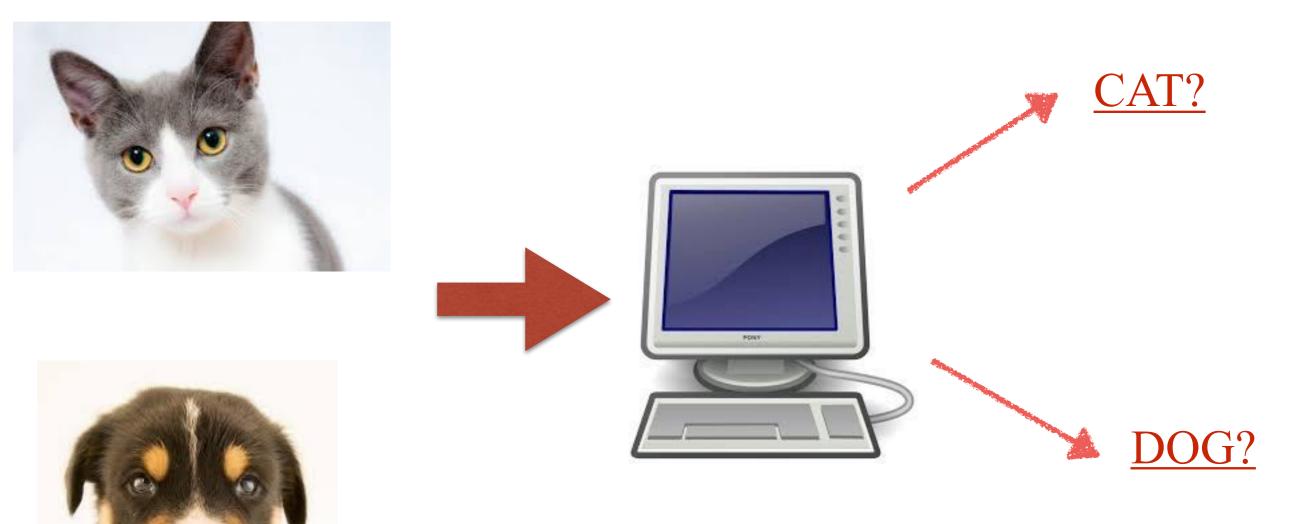
<u>Berta Margalef-Bentabol</u>, <u>Fernando Caro</u>, <u>Christoph Lee</u>, <u>Alexandre Boucau</u>d, Avishai Dekel, Joel Primack, Tom Charnock, Annalisa Pillepich, Vicente Rodriguez (+ TNG team), Emille Ishida (+ COIN initiative), <u>Helena Dominguez-Sanchez</u> ...

#### AAS 233rd meeting, Machine Learning in Astronomical Data Analysis, Seattle 2019

# AI FEVER?



## BEFORE 2012...



## TRIVIAL HUMAN TASKS REMAINED CHALLENGING FOR COMPUTERS

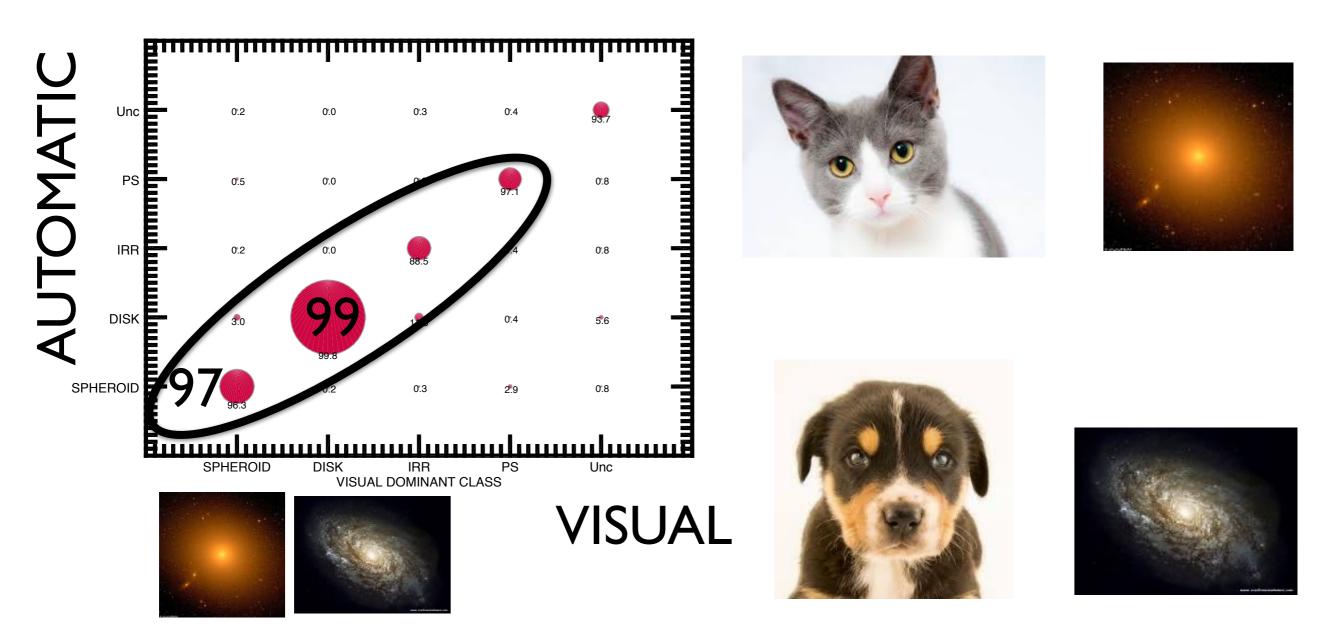
# AFTER 2012

mite	container ship	motor scooter	leopard
mite	container ship	motor scooter	leopard
black widow	lifeboat	go-kart	jaguar
cockroach	amphibian	moped	cheetah
tick	fireboat	bumper car	snow leopard
starfish	drilling platform	golfcart	Egyptian cat
grille	mushroom	cherry	Madagascar cat
convertible	agaric	dalmatian	squirrel monkey
grille	mushroom	grape	spider monkey
pickup	jelly fungus	elderberry	titi
beach wagon		ffordshire bullterrier	indri
fire engine	dead-man's-fingers	currant	howler monkey



## **IT HAS BECOME TRIVIAL...**

#### **"OUR CATS AND DOGS": GALAXY MORPHOLOGY**



**CNNs** 

DEEP LEARNING SOLVES THE PROBLEM OF GALAXY MORPHOLOGICAL CLASSIFICATION?

MHC+15b

#### TALK ON WEDNESDAY AFTERNOON -

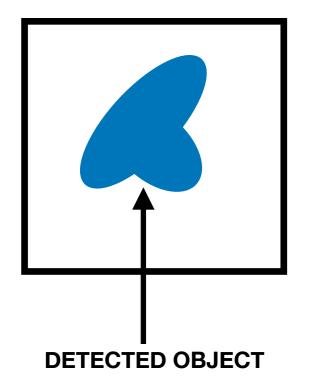
#### SPECIAL SESSION ON "MODERN" GALAXY MORPHOLOGIES

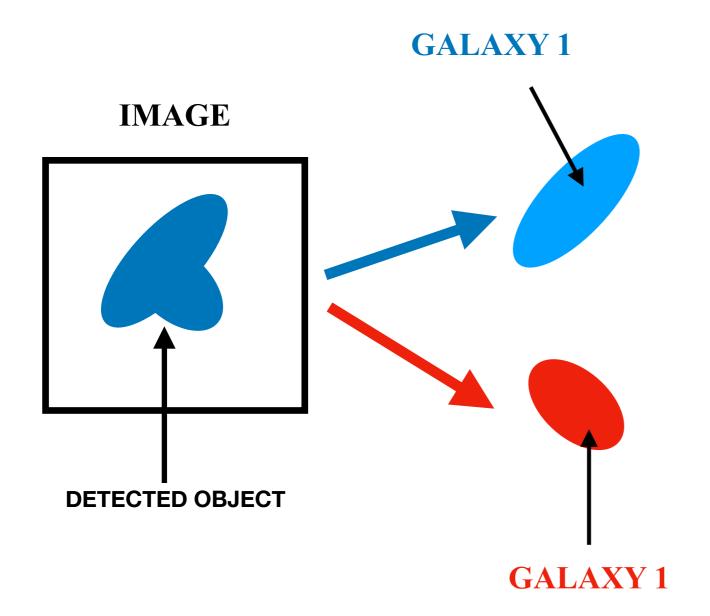
2 TAKE HOME MESSAGES FOR TODAY

#### 1. MOST OF THE PROCESSING WE DO ON IMAGES CAN BE DONE WITH AI - POSSIBLY MORE EFFICIENTLY AND MORE ACCURATELY

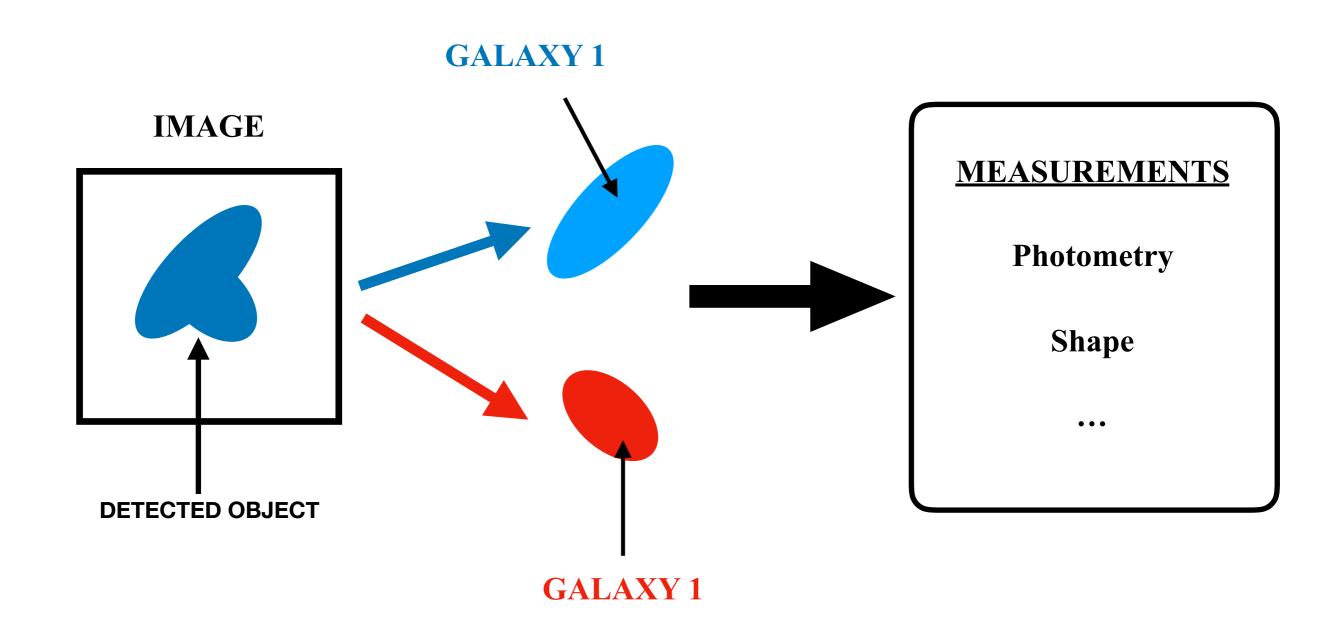
#### 2. WE CAN LEARN SOME PHYSICS BY USING AI TO LINK SIMULATIONS AND OBSERVATIONS

#### IMAGE

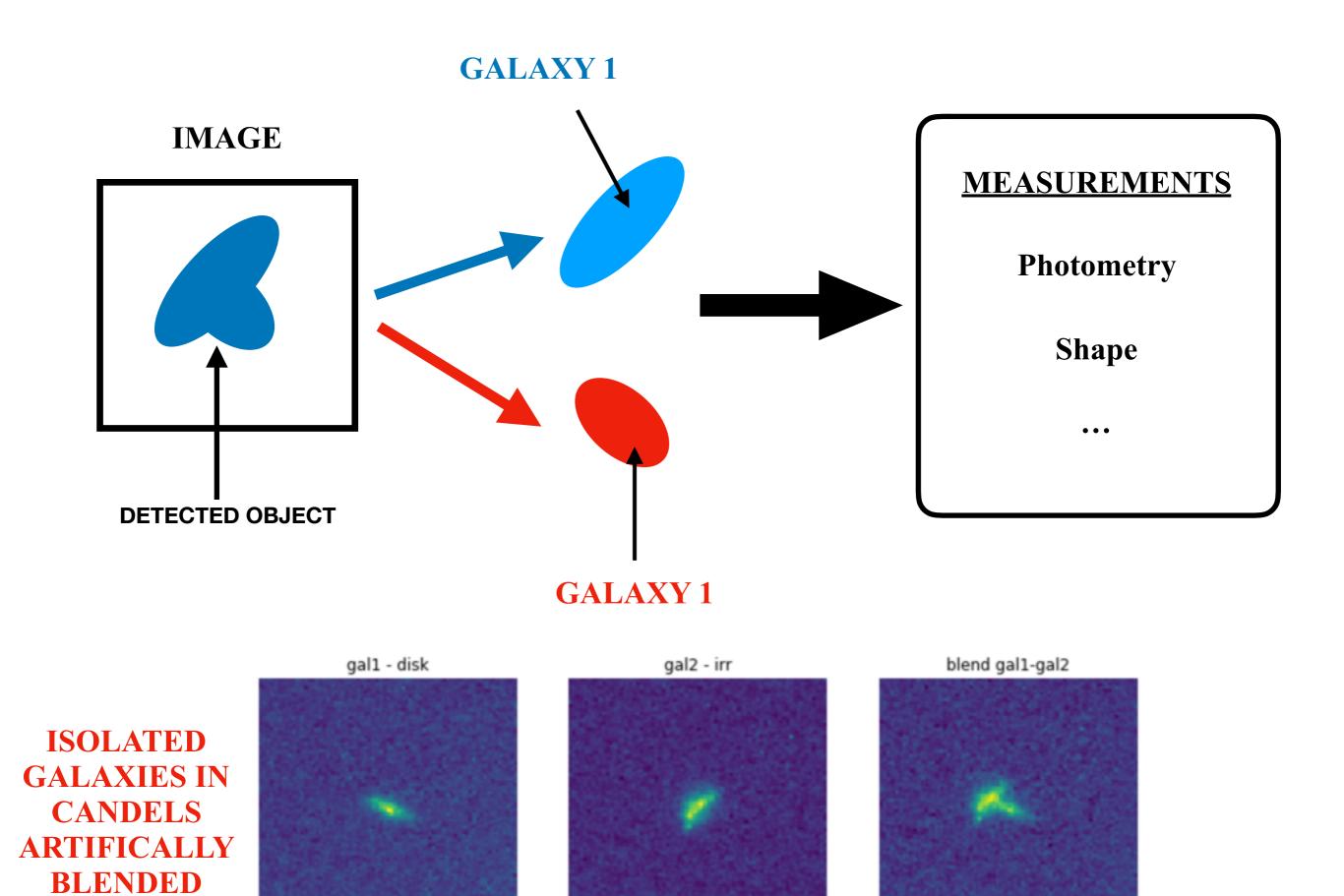




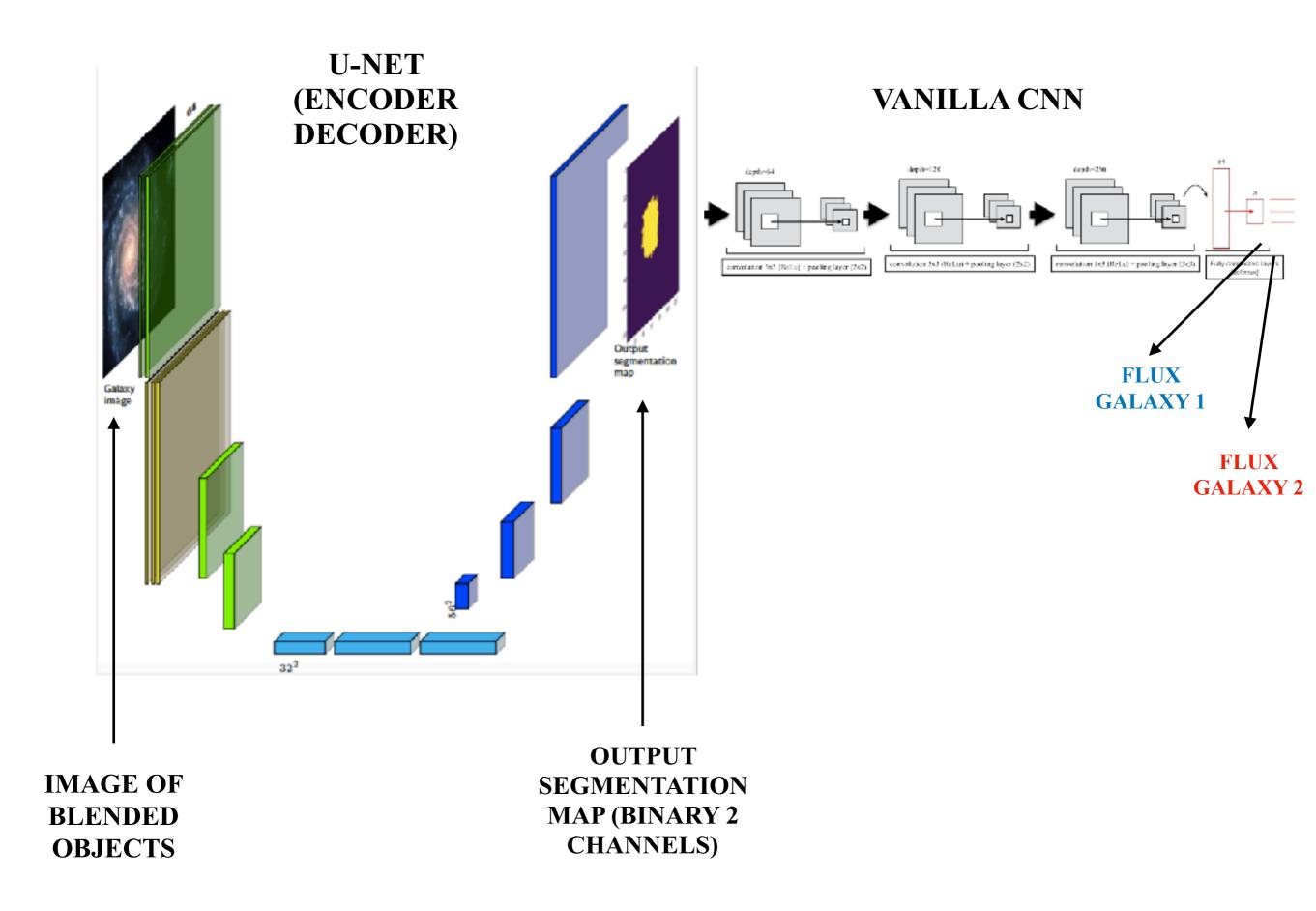
Boucaud+19



Boucaud+19

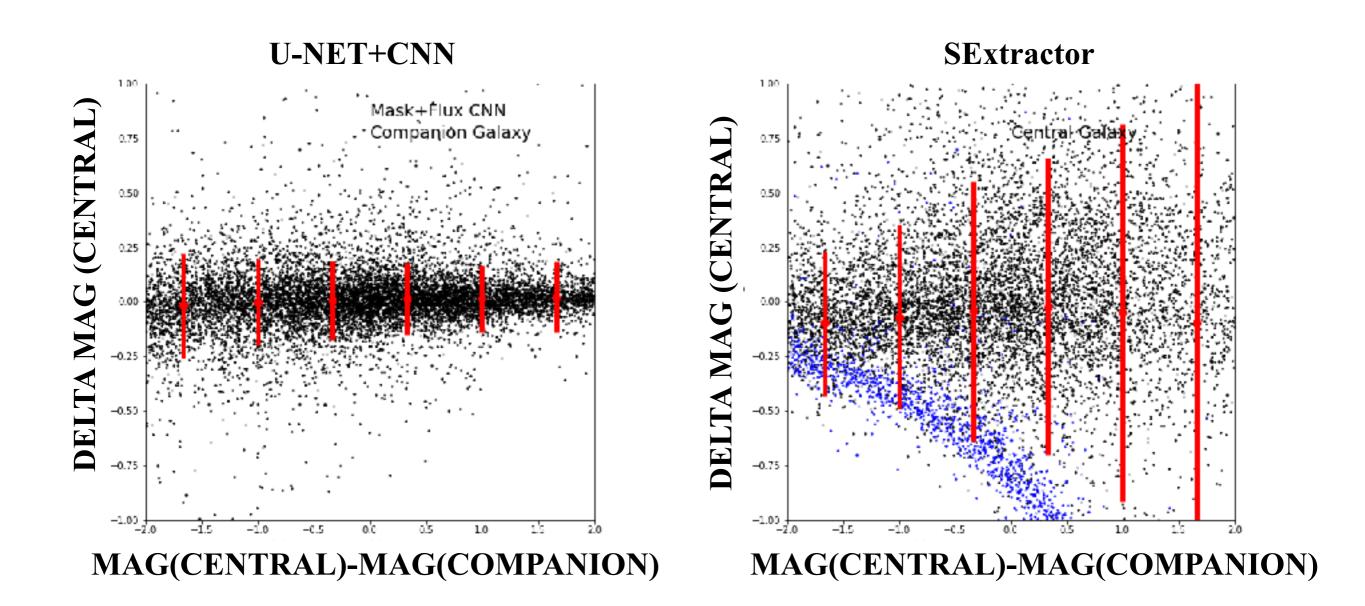


\_oucaud+19



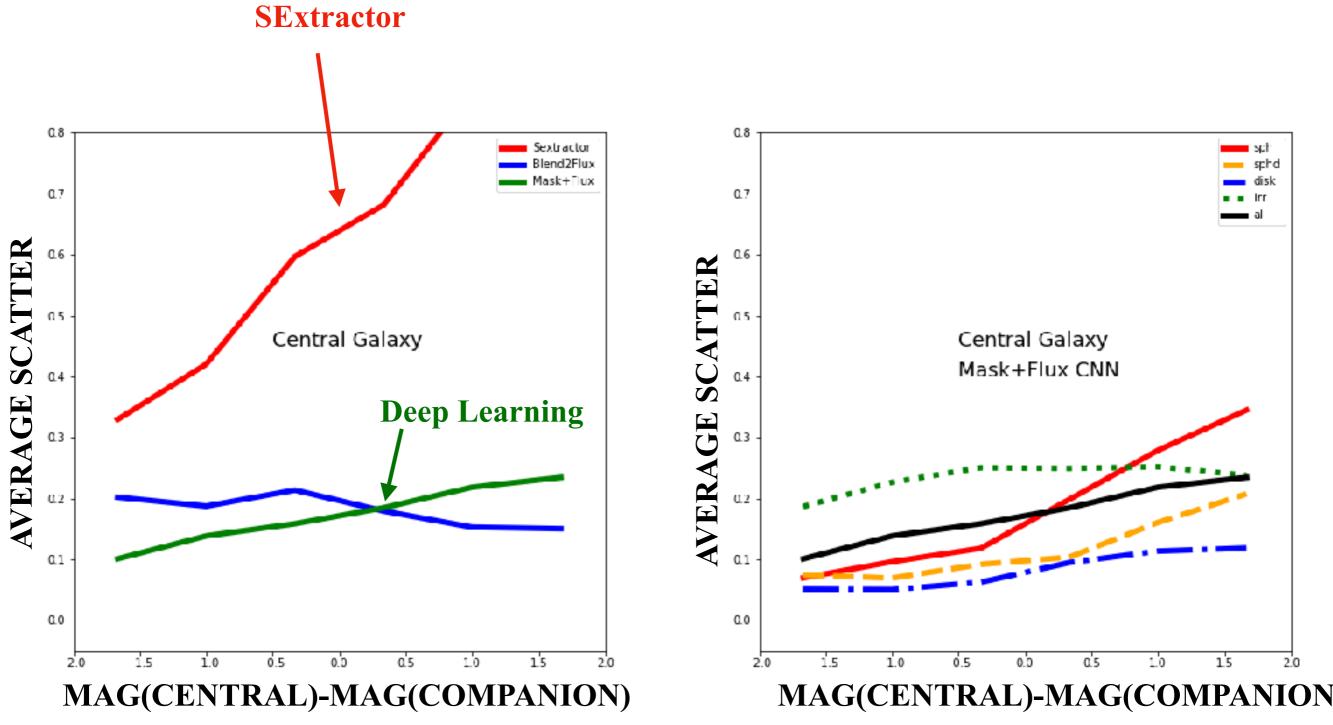
Boucaud+19

#### **PHOTOMETRY OF BLENDED SOURCES**



**Boucaud, MHC+19** 

#### **PHOTOMETRY OF BLENDED SOURCES**



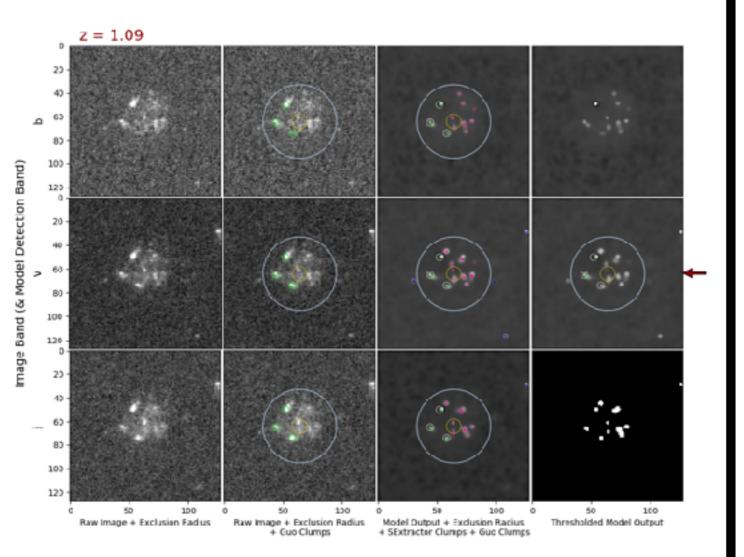
#### CONSTANT <0.2 SCATTER. A FACTOR 3-4 BETTER THAN SExtractor

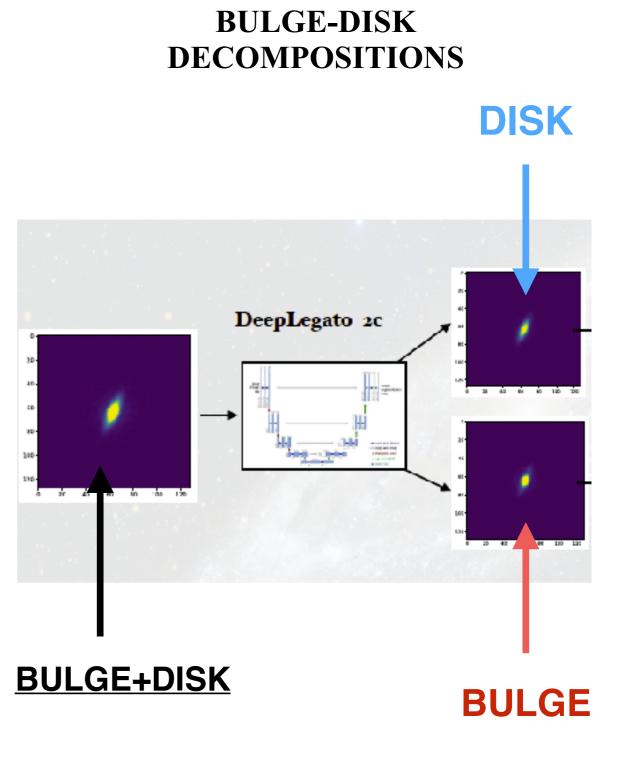
**REASONABLE BEHAVIOR FOR DIFFERENT MORPHOLOGIES** 

**Boucaud, MHC+19** 

#### **SIMILAR APPROACHES CAN BE USED FOR...**

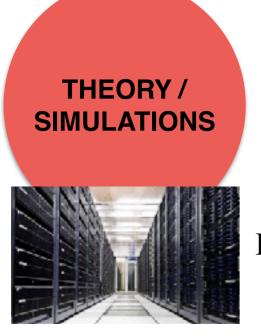
#### DETECTION OF CLUMPS IN HIGH REDSHIFT GALAXIES





Lee, MHC+19

Tuccillo, MHC+19



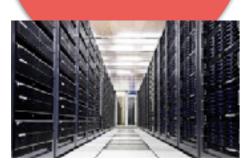
## AI TO LINK THEORY AND OBSERVATION IN THE DATA SPACE



[FULL 3D EVOLUTION HISTORY]

THEORY / SIMULATIONS

## AI TO LINK THEORY AND OBSERVATION IN THE DATA SPACE



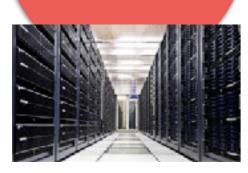
Illustris, EAGLE, Horizon-AGN ...

#### [FULL 3D EVOLUTION HISTORY]



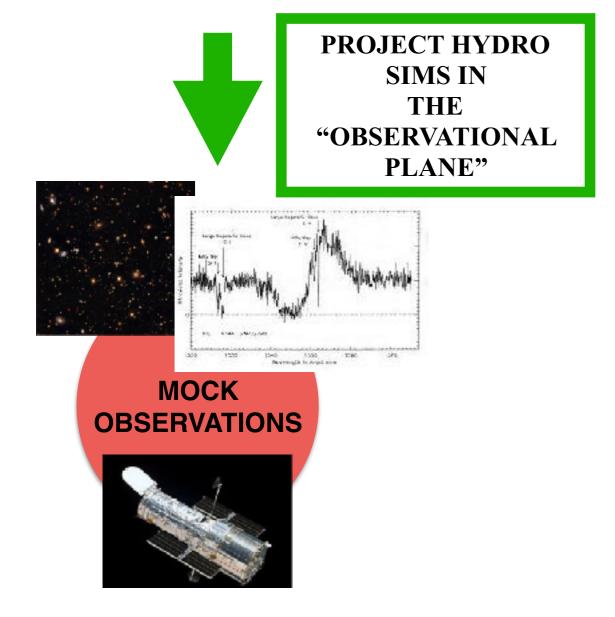
ASSUMPTIONS OF MASS TO LIGHT CONVERSION + DUST +PSF + NOISE THEORY / SIMULATIONS

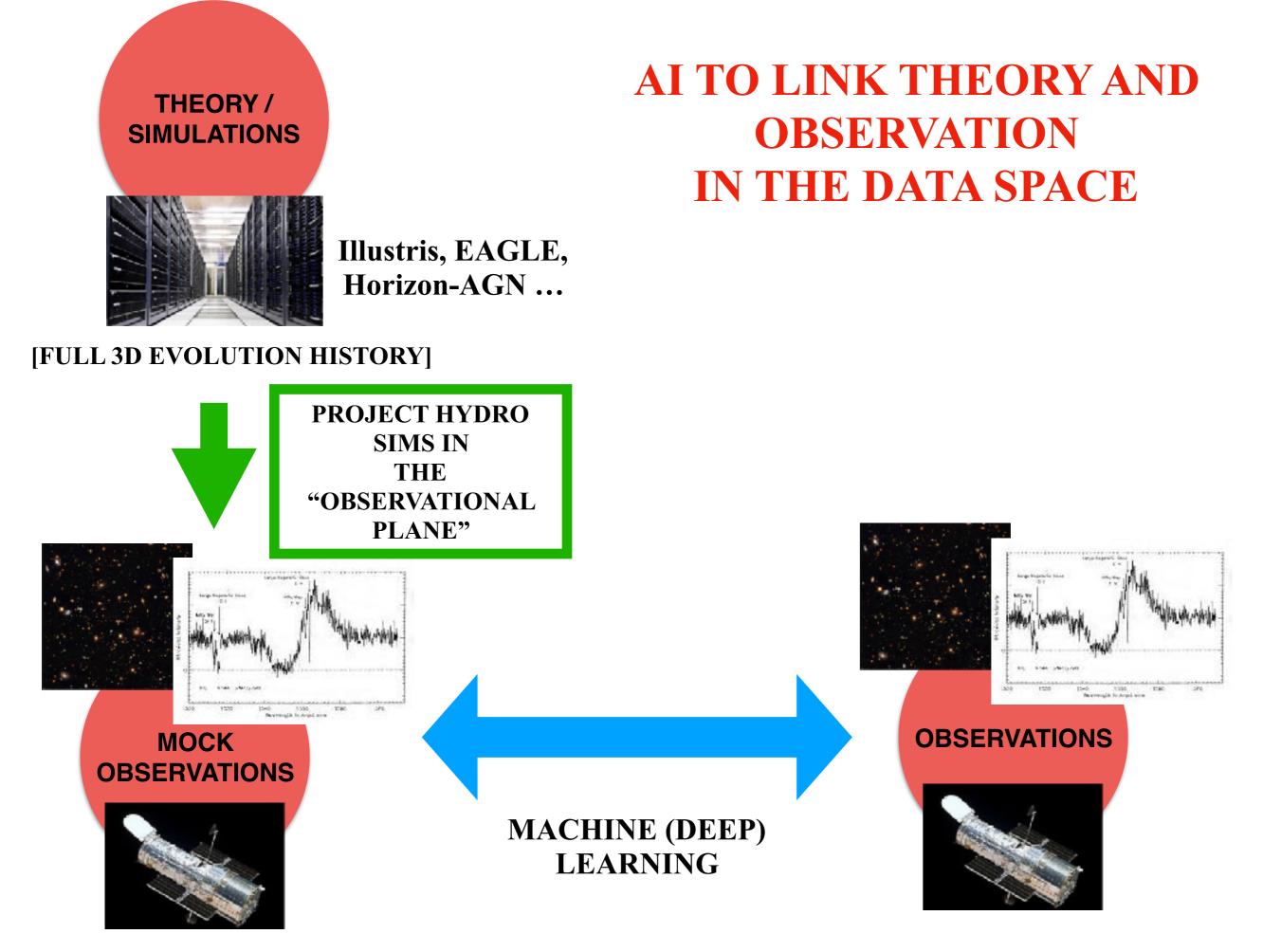
## AI TO LINK THEORY AND OBSERVATION IN THE DATA SPACE

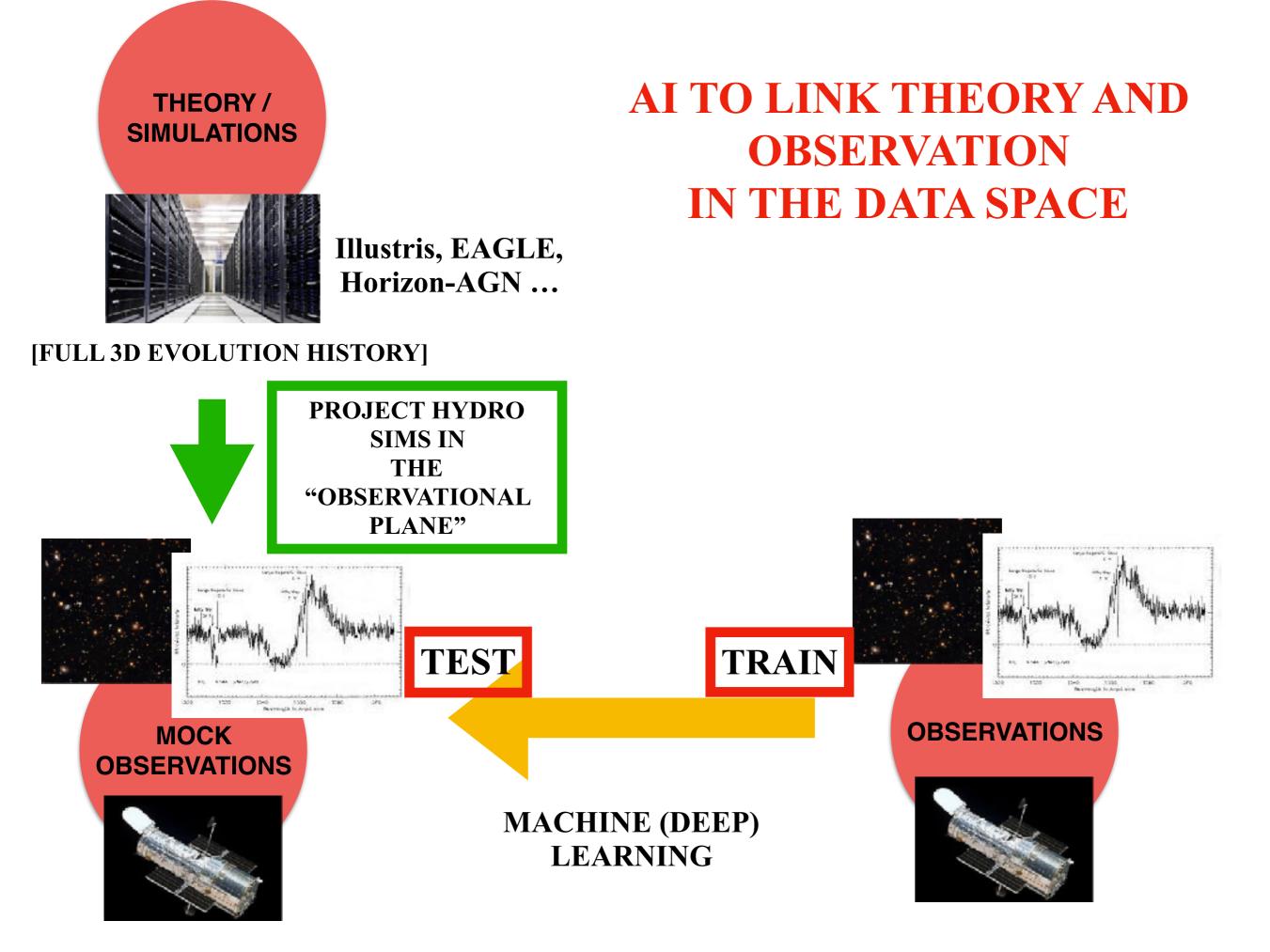


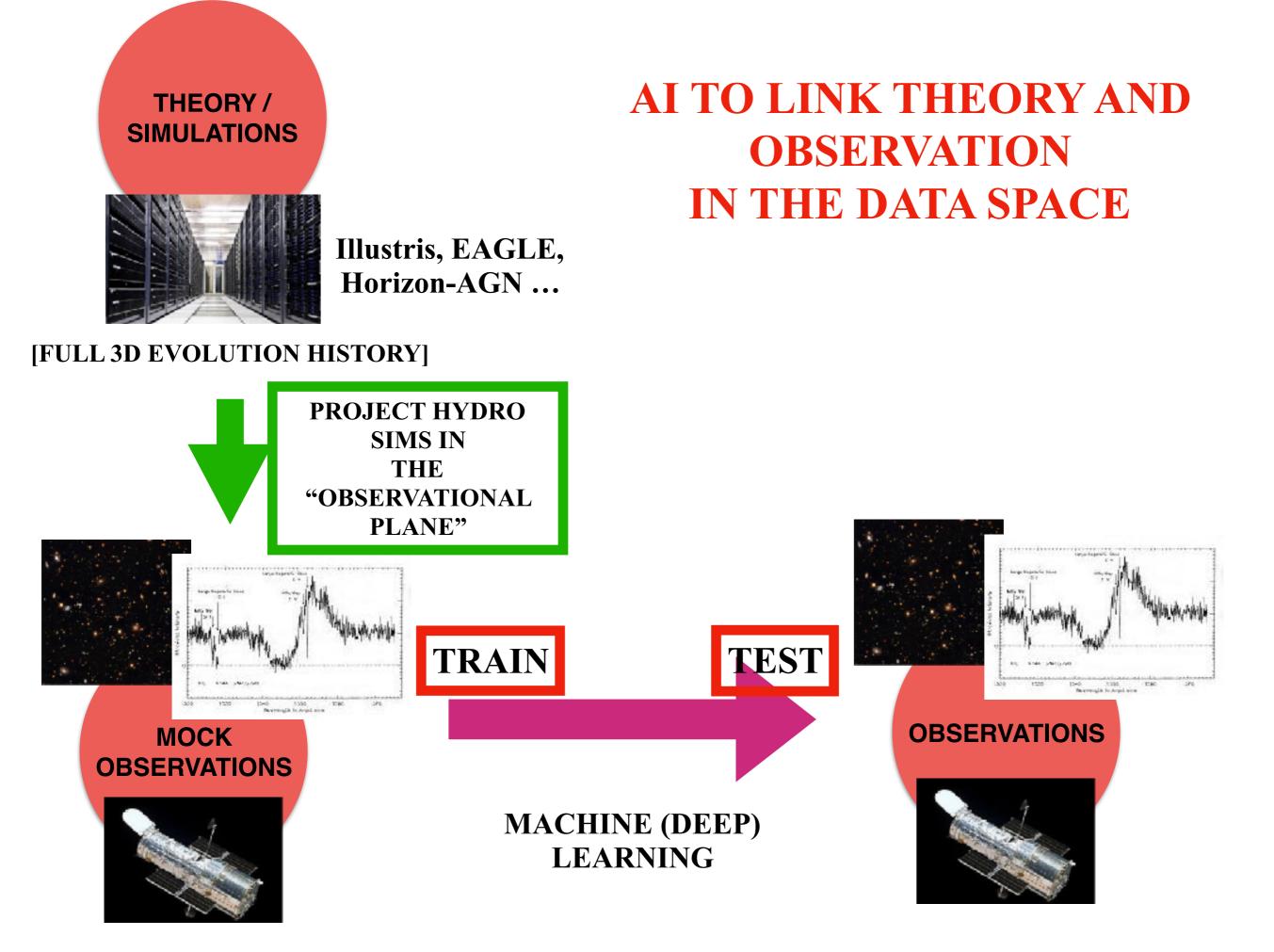
Illustris, EAGLE, Horizon-AGN ...

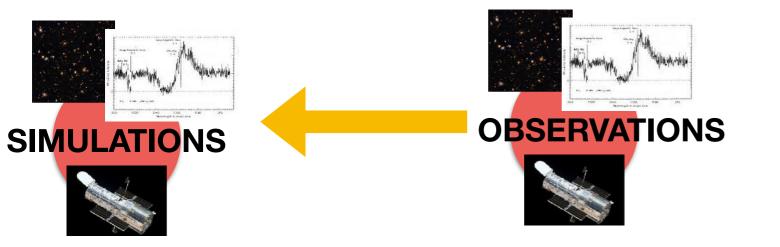
#### [FULL 3D EVOLUTION HISTORY]



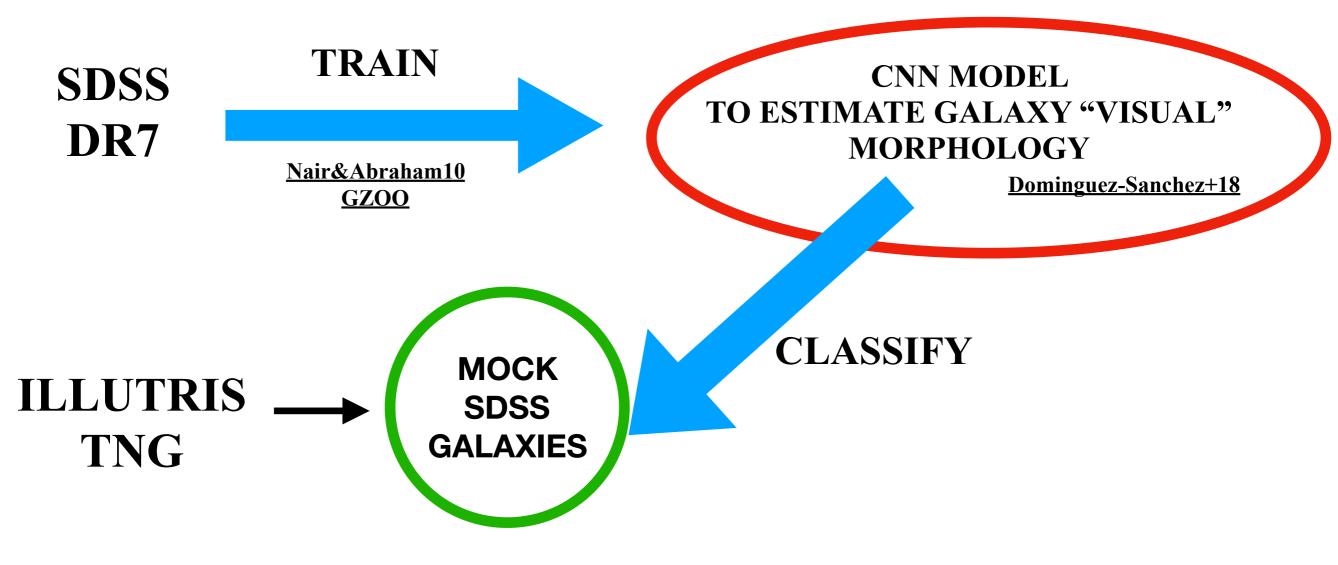












ILLUSTRIS ELLIPTICALS					
(.gj(M - /M <sub>0</sub> ) = 10.59	(Log(M+.M <sub>0</sub> ) = 10.63	Log(M+/M <sub>0</sub> ]=10.68	$Lo_{5}(M_{*},M_{\oplus}) = 10.71$		
Log(M+/M <sub>-0</sub> ) = 10.74	Log(M+,M₀) = 10.78	Leg(µ - 1µ <sub>0</sub> ] = 10.82	Log(M+JMn) = 10.87		
Log(M - ,M <sub>0</sub> ) = 10.95	Log(M-/M <sub>©</sub> ) = 11.02	$Log(\mu, i\mu_{0}) = 11.09$	1.05(M-JM_o) = 11.29		

#### **ILLUSTRIS EARLY SPIRALS**

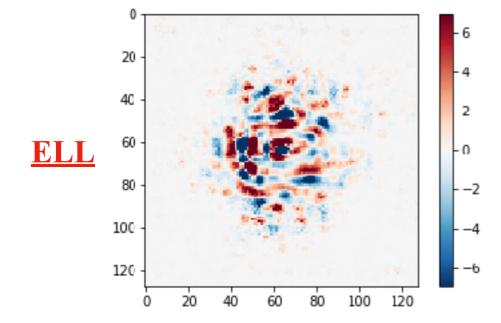
Cog(M+,M_o) = 10.69	(Log(M=,M_{0}) = 10.78	!cg;M→M <sub>3</sub> ]=10.84	$lo_{2}(M_{*},M_{\odot}) = 10.91$
Log(M+(M_a) = 10.97	Log(M+/M_0) = 11.00	Lcg!/- (µ <sub>3</sub> ] = 11.10	Log(M+,M_c) = 11.15
Log(M., M. <sub>0</sub> ) = 11.12	Log(M-/M_+1-11-39	Lcg[M(M_2]=11.37	105(M-,M_c) = 11.48
	1		•

$Log(M_{+}/M_{\odot}) = 5.54$	ILLUSTRIS SOs		$Log(M_{-}M_{\odot}) = 10.03$
(cg(M+,M <sub>0</sub> ) = 10 19	<i>Lo</i> β(M•,M <sub>Φ</sub> ) = 10.23	[tcg][M+(M <sub>2</sub> )] = 10.24	Lo <sub>2</sub> (M-,M <sub>0</sub> ) = 10.25
Log(M+.IM <sub>in</sub> ) = 10 34	Log(M-,M⊙) = 10.41	Log(M - (M <sub>0</sub> ) = 10.43	Log(M+iM_n) = 10.43
1			•
Log(M - ,M <sub>0</sub> ) = 10.47	Log(M-,M <sub>0</sub> ) - 10,48	Log[M . (M <sub>D</sub> ] = 10.56	Log(M-,M <sub>G</sub> ) = 10.56

**ILLUSTRIS LATE SPIRALS** 

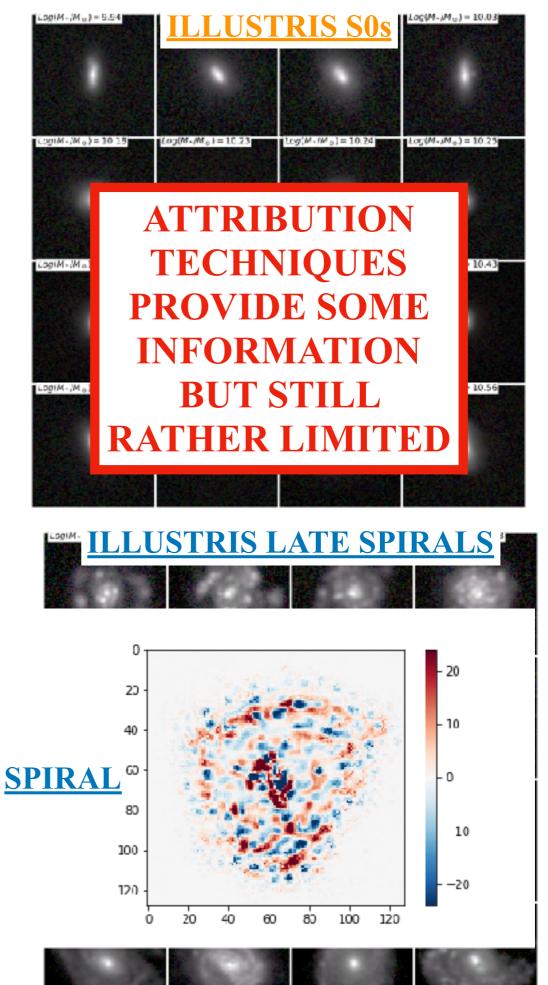
Содим.,Мо) = 10.03	Log(M.,M.,)=10.32	Lcg;M. (M <sub>2</sub> ) = 10.19	Log(M.: M.s.) = 10.25
LogiM+,(M <sub>in</sub> ) = 10 32	Log(M+.M+,) = 10.38	Lcg://- /// 3] = 10.44	Log (M+ M=) = 10.50
Lbg(M - ,M <sub>0</sub> ) = 10.55	Log(M-,M <sub>©</sub> ) = 10.64	Lcg(M-144 <sub>p</sub> ) = 10.70	Log(M-1M <sub>0</sub> ) = 10.80



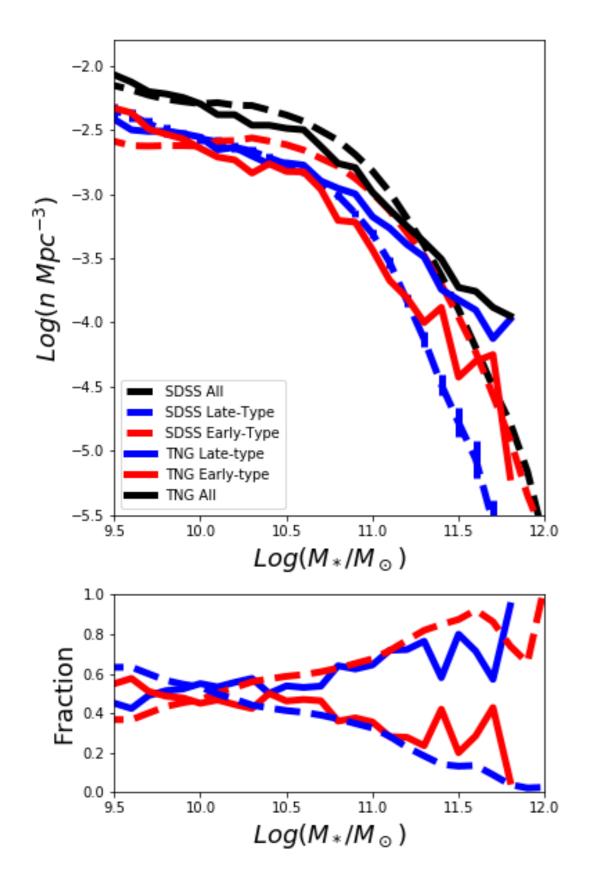


#### **ILLUSTRIS EARLY SPIRALS**

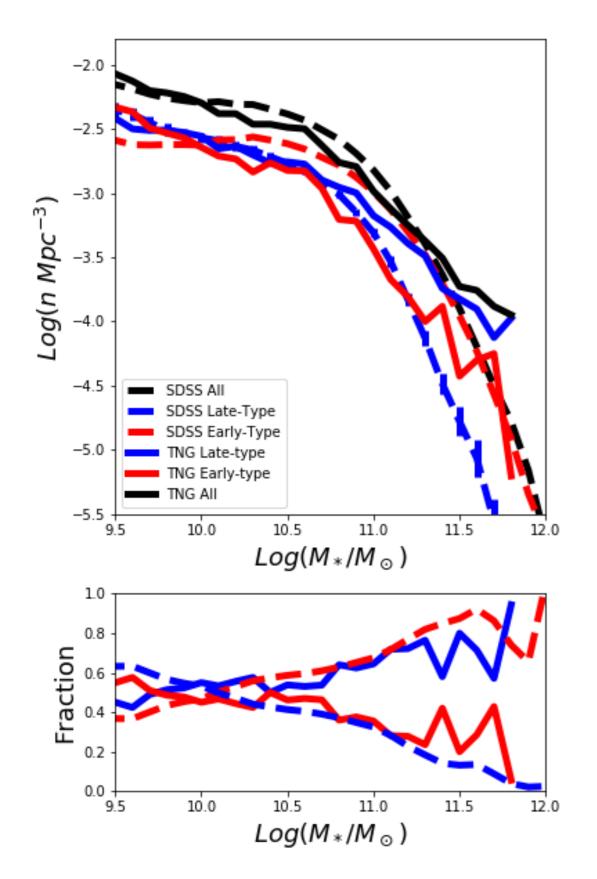
Cog(M+/M <sub>0</sub> ) = 10.63	Log(M-,M <sub>0</sub> ) = 10.78	1cg;W-(W <sub>p</sub> ]=10.84	Log(M-,M_0) = 10.91
Log(M · /M <sub>o</sub> ) = 10 97	Log(MM <sub>0</sub> ) = 11.00	Lcg!⊬-(µ <sub>0</sub> ]=11.10	Log(M-1Ma) = 11.15
Log(M-,M <sub>0</sub> ) = 11 19	Log(M-,M <sub>&amp;</sub> ) = 11.39	Lcg[M./M <sub>p</sub> ] = 11.37	Log(M-,M <sub>☉</sub> ) = 11.48

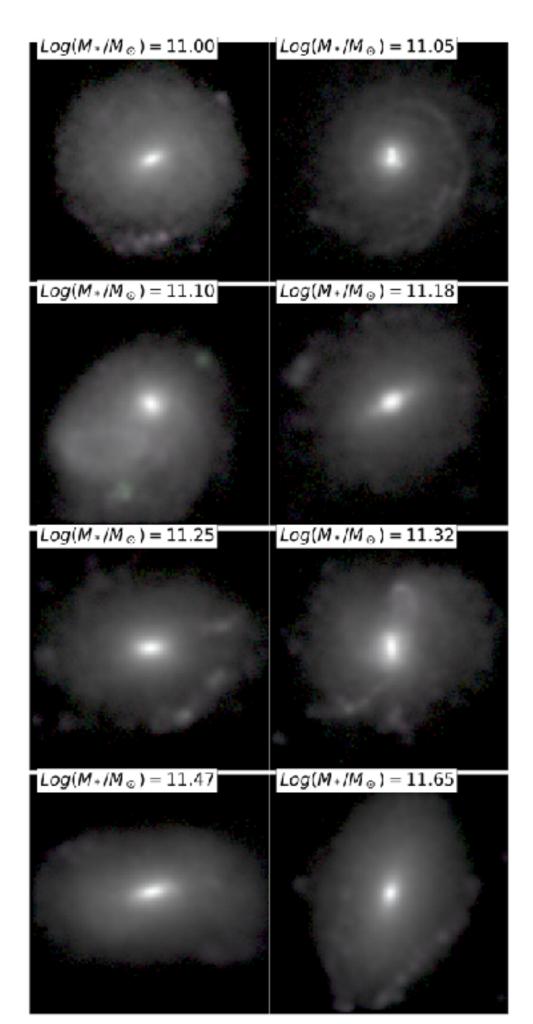


#### THE HIGH MASS END OF THE STELLAR MASS FUNCTION OF ILLUSTRIS TNG IS DOMINATED BY DISKS

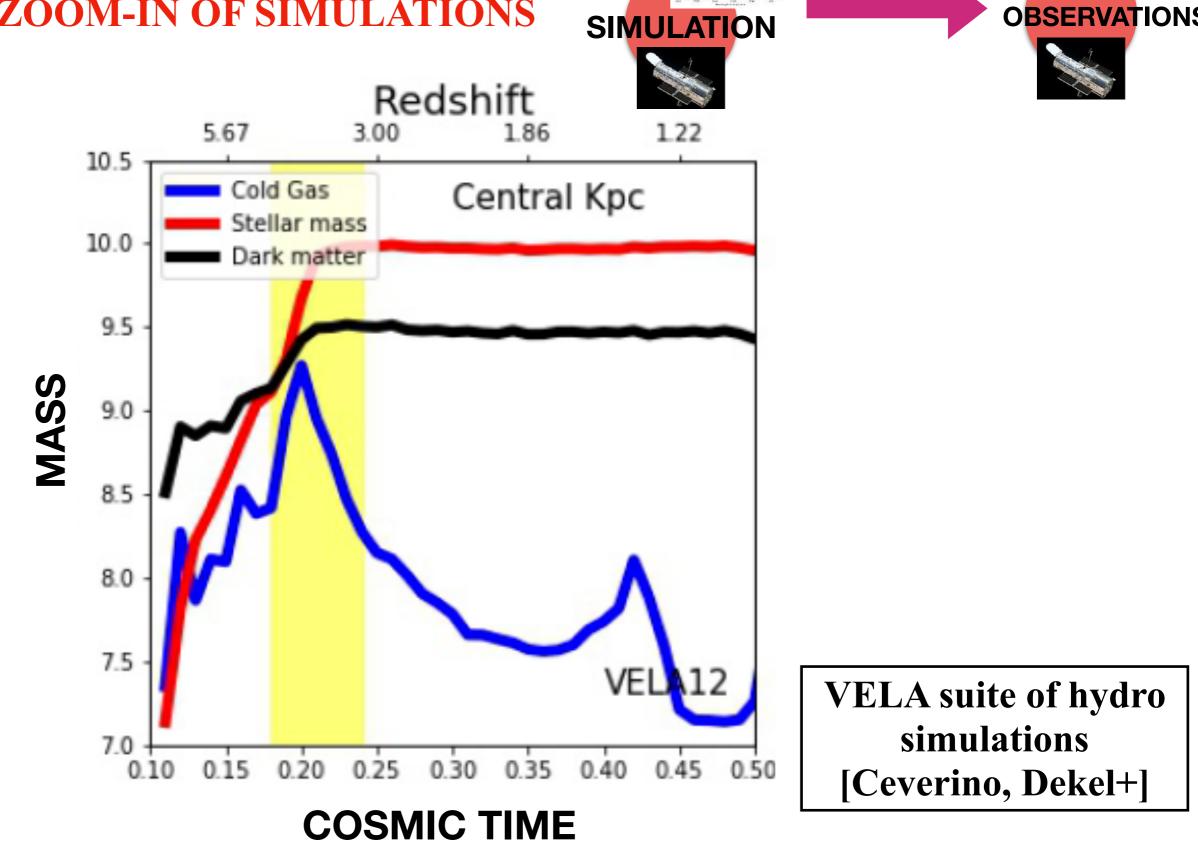


#### THE HIGH MASS END OF THE STELLAR MASS FUNCTION OF ILLUSTRIS TNG IS DOMINATED BY DISKS



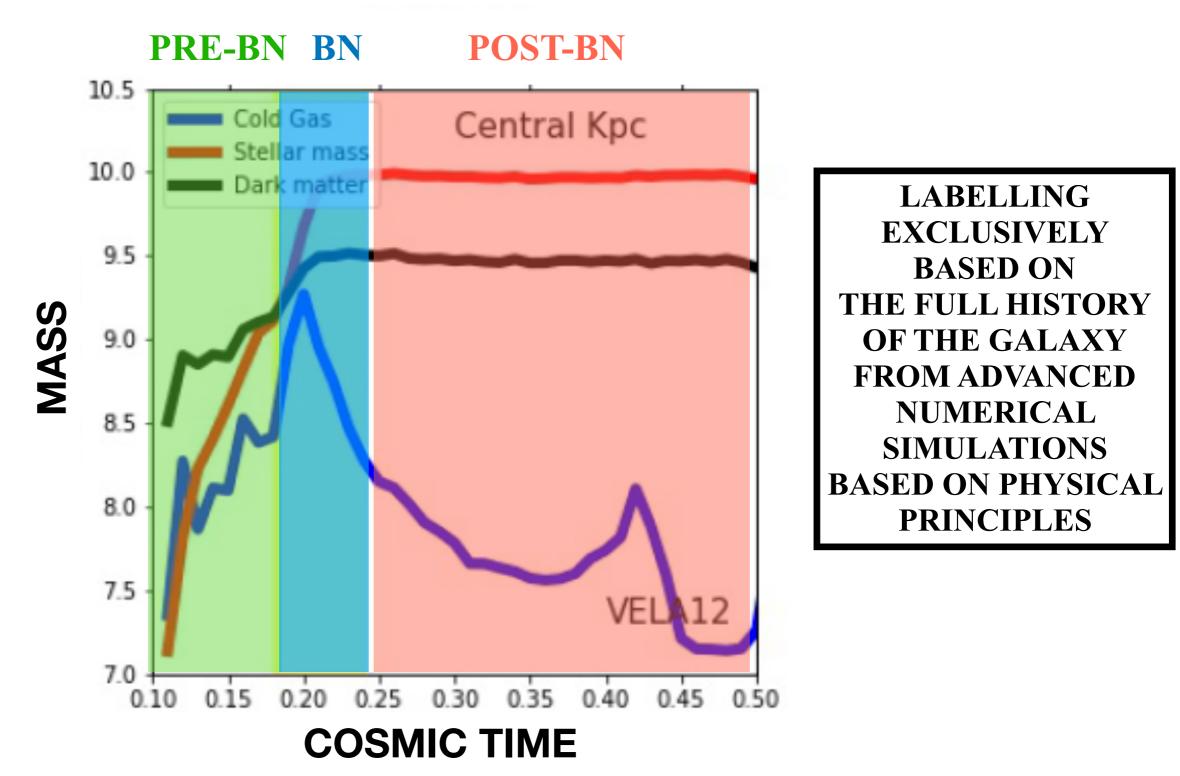


## IDENTIFYING COMPACTION IN THE VELA ZOOM-IN OF SIMULATIONS



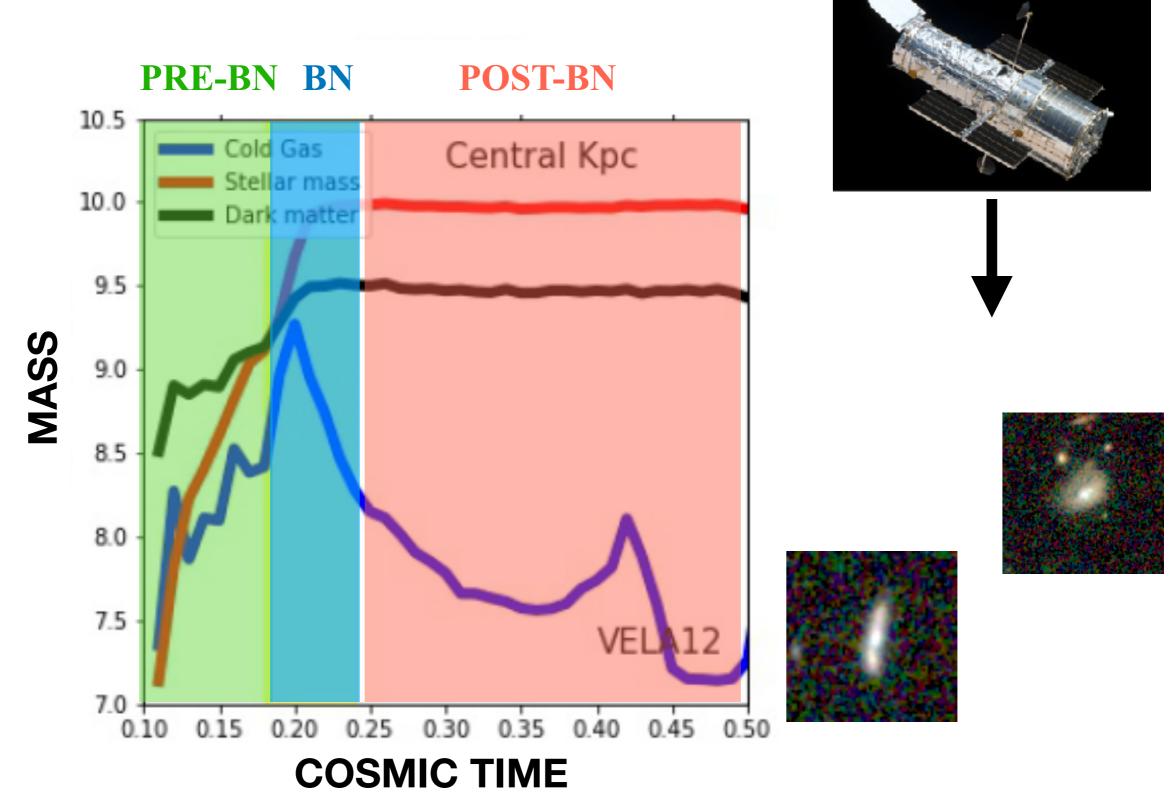
Dekel&Burkert+14, Ceverino+15, Zolotov+15, Tacchella+16a,b, Dekel+18

#### IDENTIFYING COMPACTION IN THE VELA ZOOM-IN OF SIMULATIONS



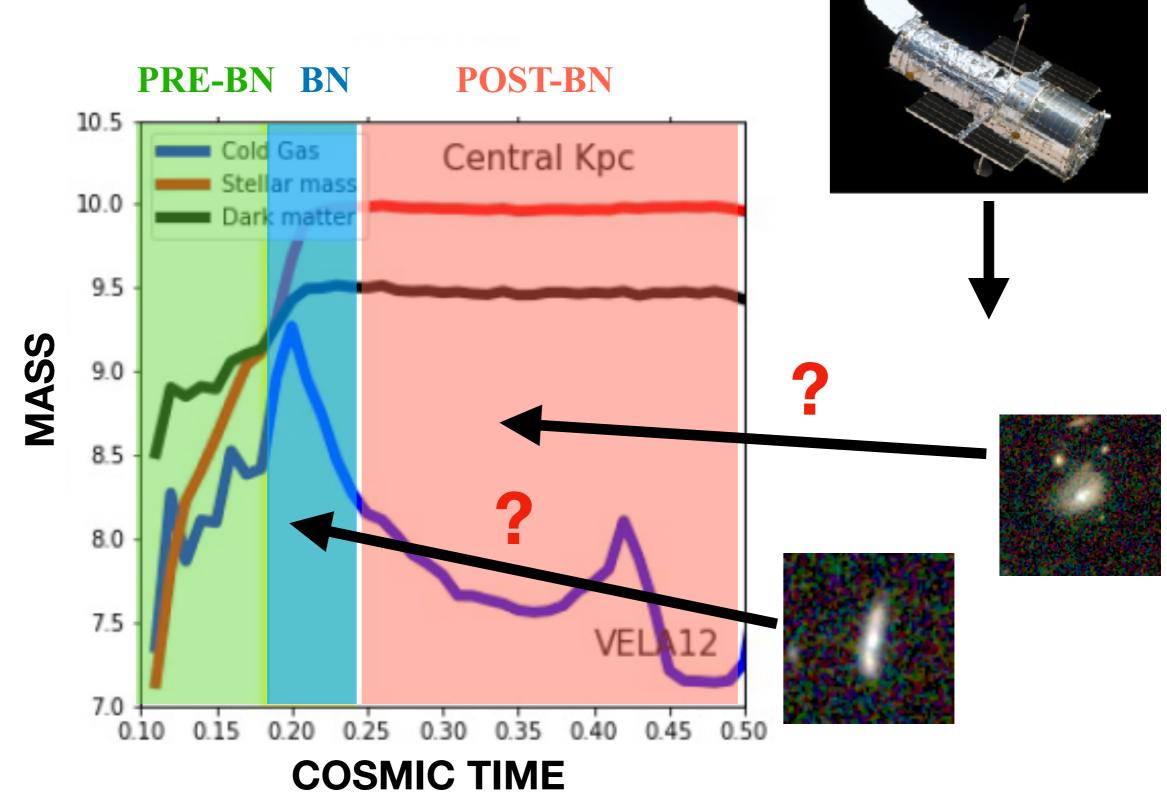
Dekel&Burkert+14, Ceverino+15, Zolotov+15, Tacchella+16a,b, Dekel+18

#### IN THE OBSERVATIONS WE ONLY HAVE ONE SNAPSHOT OF A GALAXY AT A GIVEN TIME

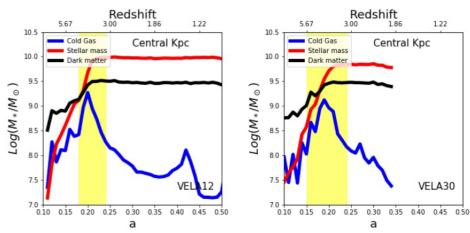


Dekel&Burkert+14, Ceverino+15, Zolotov+15, Tacchella+16a,b, Dekel+18

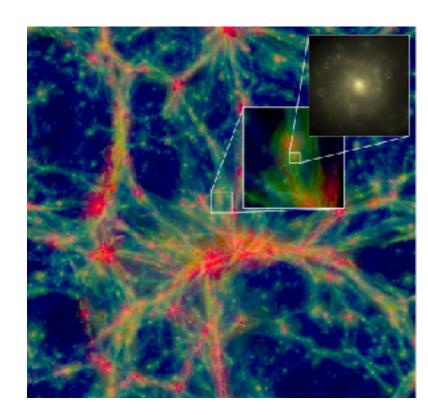
#### HOW CAN WE ESTIMATE THE PHASE FROM A UNIQUE IMAGE?

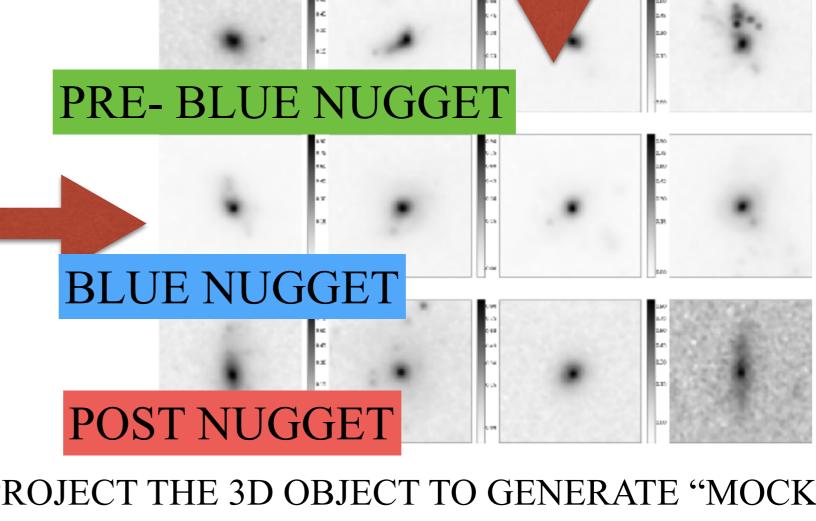


Dekel&Burkert+14, Ceverino+15, Zolotov+15, Tacchella+16a,b, Dekel+18



## USE THE FORMATION HISTORY OF EACH GALAXY TO LABEL IMAGES ...

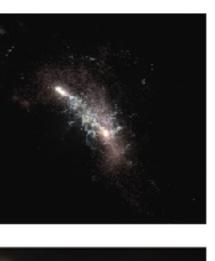




[COSMOLOGICAL SIMULATION]

PROJECT THE 3D OBJECT TO GENERATE "MOCK" IMAGES AS OBSERVED BY HUBBLE

#### Pre-Blue-Nugget-Stage





Blue-Nugget-Stage



#### **VELA HIGH RESOLUTION**

#### **VELA HST RESOLUTION**



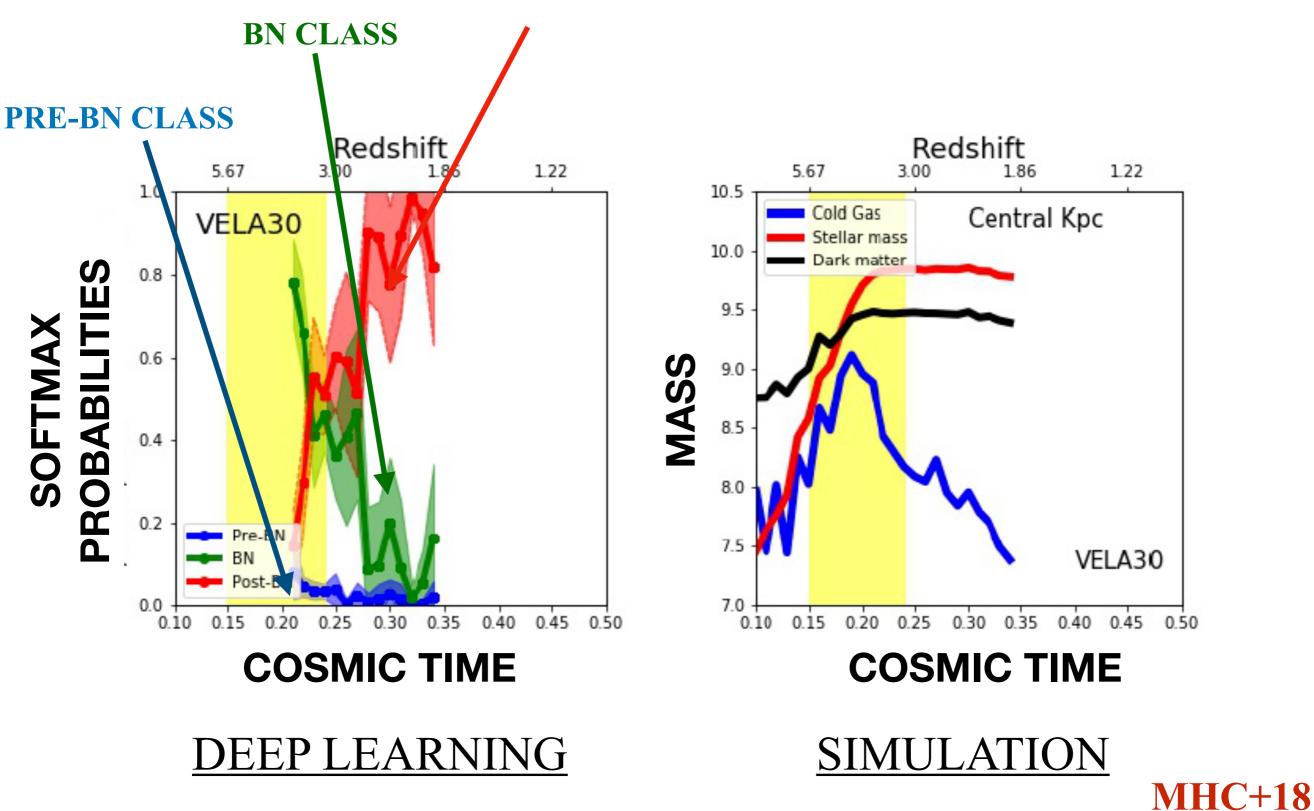




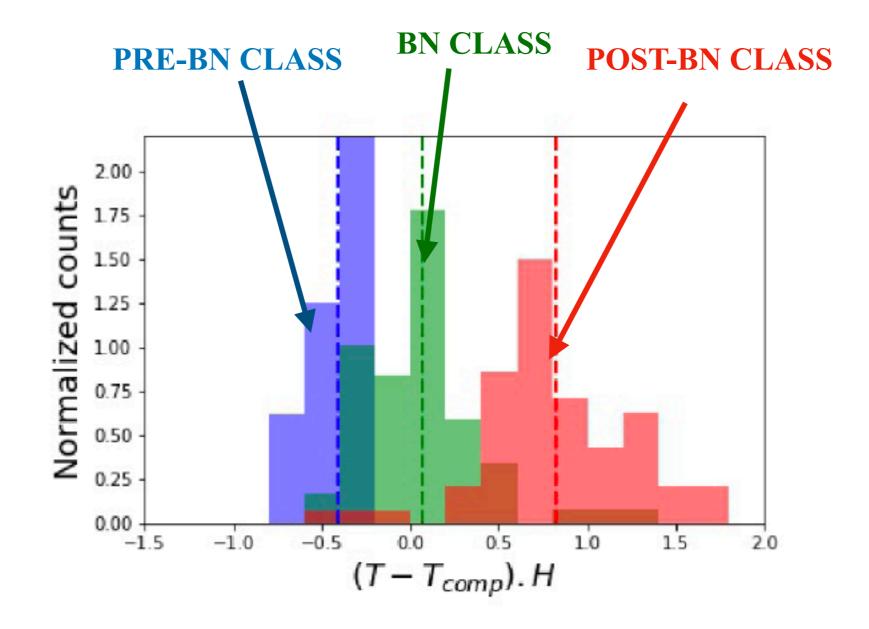
#### **CANDELS**

# THE OUPUT SOFTMAX PROBABILITY TRACKS THE GAS MASS

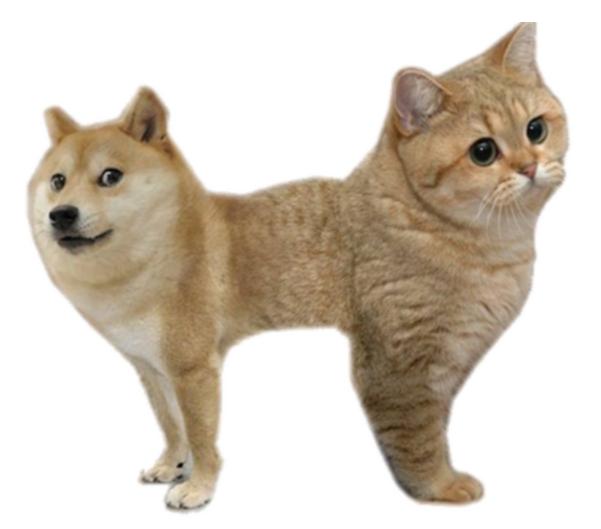




### **CONSTRAINTS ON THE OBSERVABILITY TIMESCALE**



## **CAPTURING THE MODEL UNCERTAINTY**



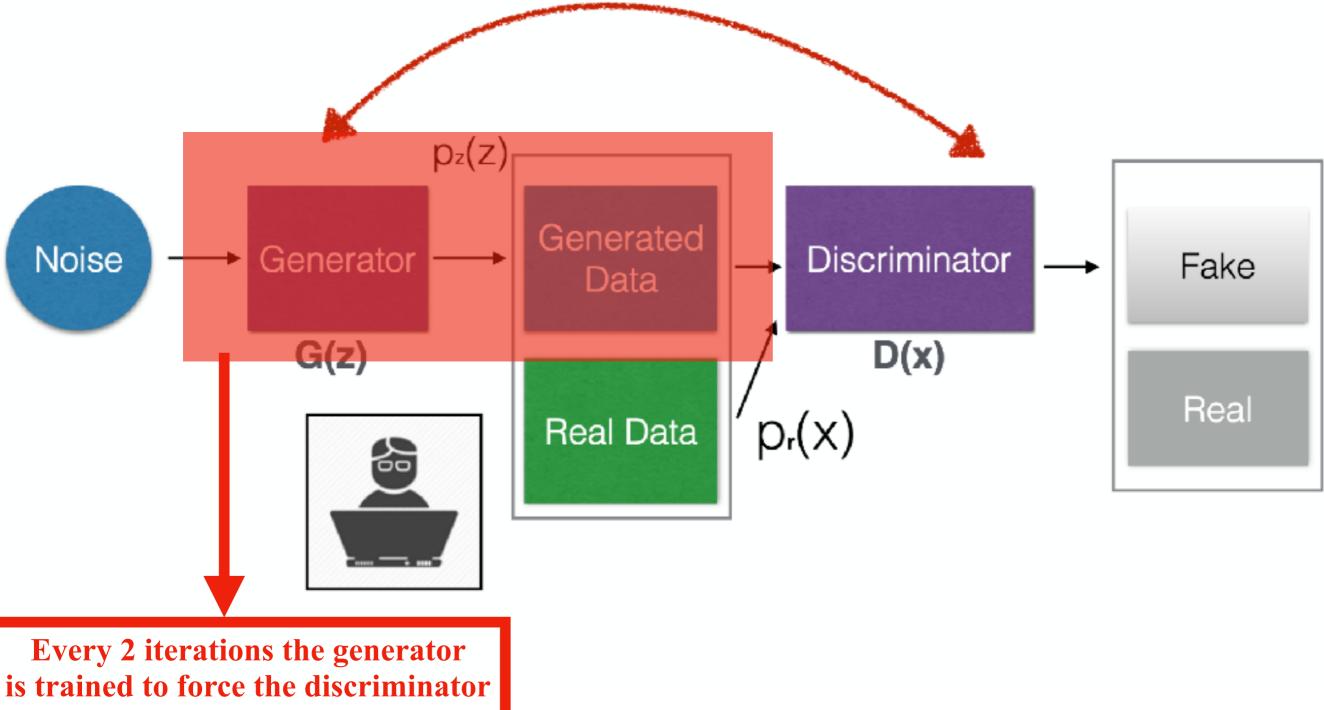
#### SEVERAL OPTIONS EXIST IN THE LITTERATURE...

#### **NO TIME TO TALK ABOUT BNNs**

## **GENERATIVE ADVERSARIAL NETWORKS**

#### (Goodfellow+)

#### **TWO COMPETING NETWORKS**

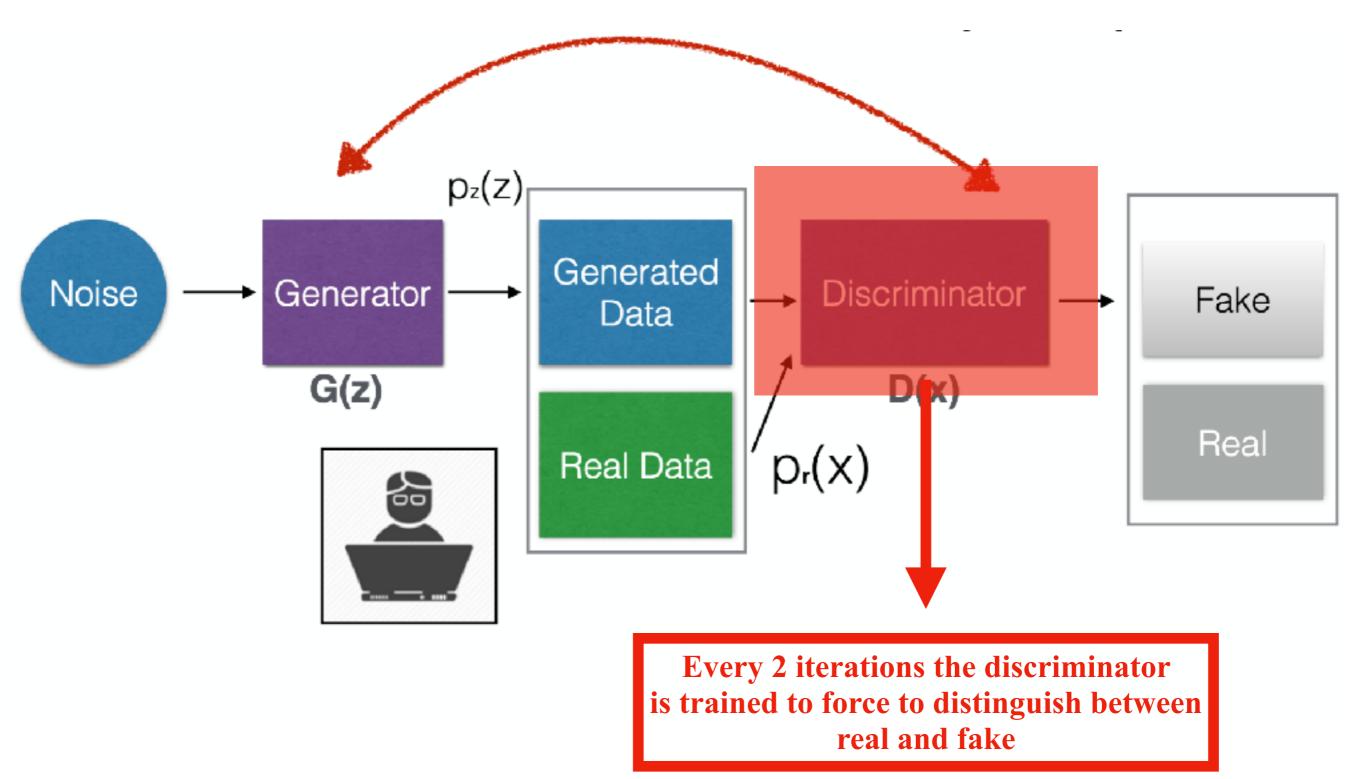


to classify as real

## **GENERATIVE ADVERSARIAL NETWORKS**

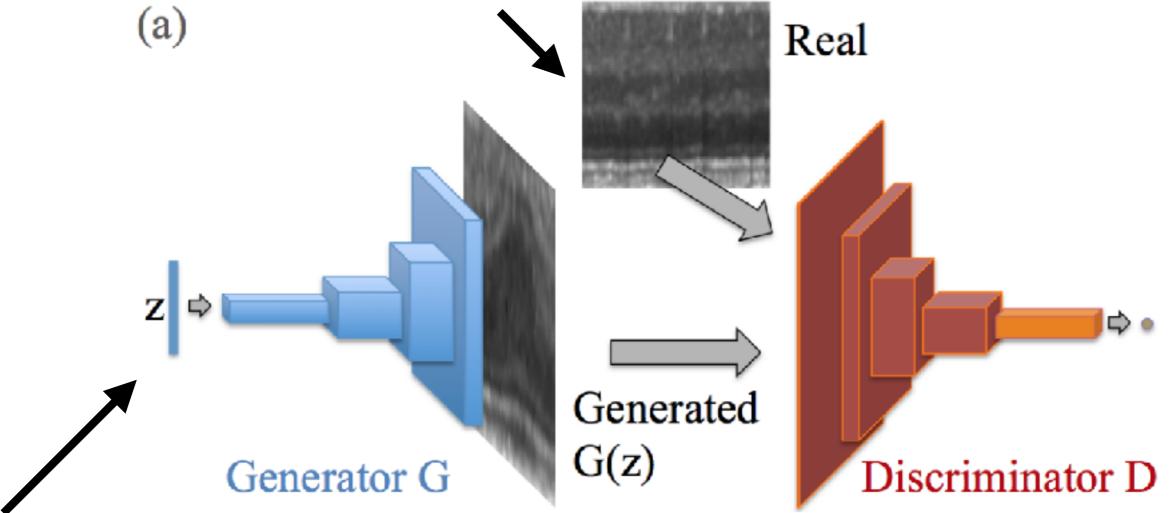
#### (Goodfellow+)

#### **TWO COMPETING NETWORKS**



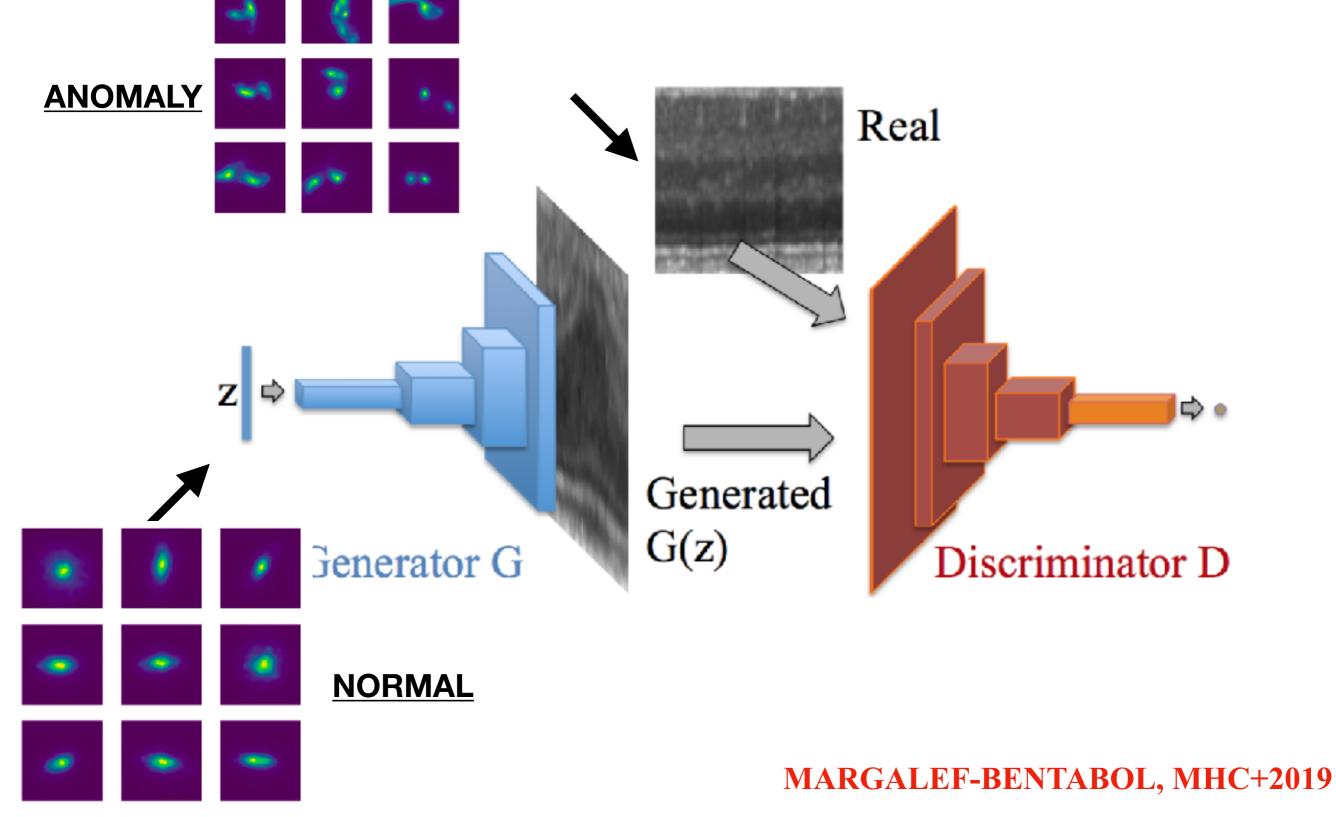
# ANOMALY DETECTION WITH GANs

## **OBSERVATIONS**

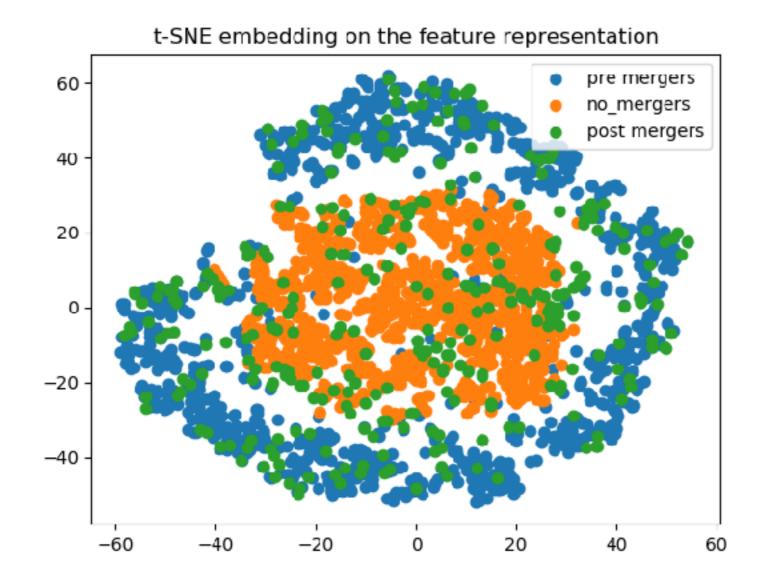




## A PROOF-OF-CONCEPT CASE: UNSUPERVISED DETECTION OF MERGERS WITH GANs

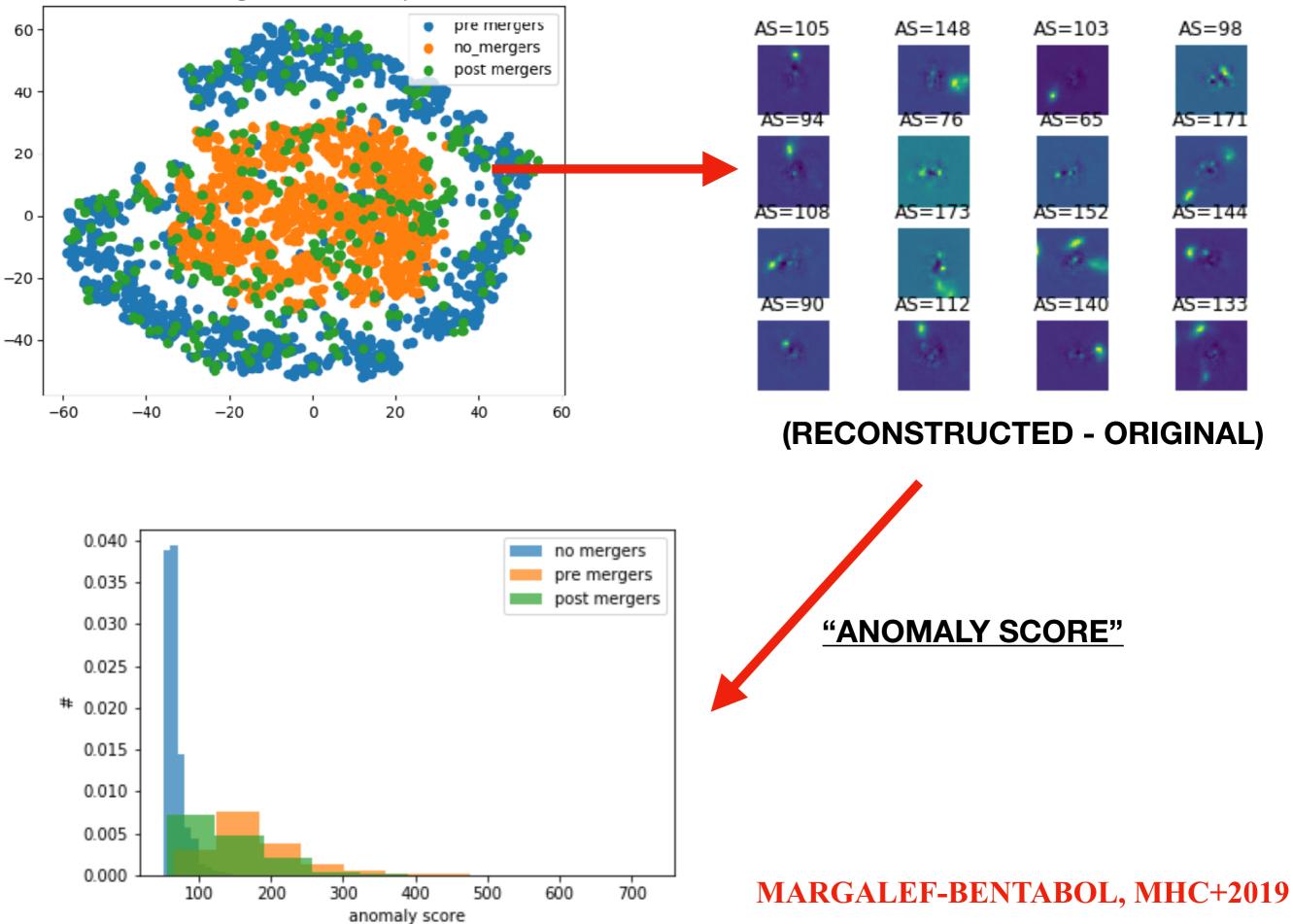


#### **MERGERS OCCUPY A DIFFERENT REGION IN THE GAN GENERATED LATENT SPACE**



#### **MARGALEF-BENTABOL, MHC+2019**

t-SNE embedding on the feature representation



2 TAKE HOME PROVOCATIVE ? MESSAGES FOR TODAY

#### 1. MOST OF THE PROCESSING WE DO ON IMAGES CAN BE DONE WITH AI - POSSIBLY MORE EFFICIENTLY AND MORE ACCURATELY

**DETECTION, PHOTOMETRY, PHOTOZ's, MORPHOMETRY ...** 

#### 2. WE CAN LEARN SOME PHYSICS BY USING AI TO LINK SIMULATIONS AND OBSERVATIONS

WITH SUPERVISED ML GOING BACK AND FORTH FROM SIMS TO DATA WITH UN-SUPERVISED ML TO MEASURE SIMILARITIES